

Astronomy ESFRI & Research Infrastructure Cluster ASTERICS - 653477



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# DNN classification of signals and glitches in time-domain gravitational-wave data

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### Gravitational-wave astronomy



5 binary black hole mergers and 1 binary neutron star merger detected so far!

## **Glitches representation**

- Studies to apply machine learning to the problem of the glitch identification are mainly based on spectrograms (GravitySpy, DeepLearning,..)
  - ✓ Deep-learning performs well on images
  - Disadvantages:
    - Volume of data (big images)
    - Spectrogram parameters/choice dependent
    - Risk of loosing information due to manipulation
    - Deep learning algorithms learn on raw data
- Time series representation
  - ✓ full information
  - Reduced volume of data



Time (s

## **General ideas**

- Study, identify and reduce the <u>transient noise</u> present in the gravitational wave detectors through <u>deep learning techniques</u>
  - Raw <u>time-series</u> as input instead of frequency-time representations (spectrograms)
  - Both strain data and auxiliary channels
  - Try different kind of deep-learning algorithms
- Final goal: analyse single-detector data

## **Single-detector time**

- Transient noise (behaviour of the instruments or complex interactions between the instruments and their environment)
  - ✓ "instrumental glitches", non-Gaussian short duration artefacts
  - ✓ mimic the gravitational wave signal.
- Current pipelines: signal has to appear in coincidence in two or more detectors
  - $\checkmark$  distinguish true astrophysical signals from the transient noise
  - highly reduces the number of false positives allowing to detect gravitational waves with very high statistical confidence.

#### Single-detector time marginally exploited

2.7 months in O1+O2 => could contain 3 events



### Glitches, noises and signals







## **Current activity**

- First step: prepare samples for the training and test
  - Training on the basis of the strain morphology

#### Generator of 3 classes of events only:

- Detector noise without loud glitches (Gaussian-noise)
  - Taken from real data when nor glitches nor signals are present
- Gaussian-noise + glitches
  - Glitches occurring times taken from cWB analysis
- Gaussian-noise + astrophysical signals
  - Signals = BBH with randomised parameter

Generator able to produce each of the 3 classes selecting randomly a piece of random noise and, if needed, adding randomly glitches or signals

 $\checkmark$  Data whitened and accompanied by the PSD to calculate the SNR

### Ongoing / next steps

- ★ setup a 1D Convolutional Neural Network (CNN) to distinguish the 3 classes of events,
- try other algorithms: recurrent neural networks (RNN), Long-Short Term Memory (LSTM),
- ★ Add features:
  - \* environmental channels (multi-instance learning),
  - \* more complicated signal/glitch morphologies,
  - \* study causality (not only corelation) between channels,
  - \* Compression to decrease the size of DNN (e.g. Bayesian compression, arXiv:1705.08665).

# RNN - LSTM

#### Neural networks for processing sequential data

- Keep a summary of the past sequence in their memory or so-called hidden state, which is updated whenever a new input token is presented.
- LSTMs incorporate a gating mechanism which controls to what extent the new input is stored in memory and the old memory is forgotten.



## **Environmental channels**

Hundreds of thousands auxiliary data streams, auxiliary channels, monitors status of the detector (e.g, state of the control loops) and of its physical environment.

Now: correlation-based techniques used to identify the coupling of a noise source with an observed disturbance

- ✓ UPV and Excavator: based on time coincidence only (Virgo)
- Weak points: need many (100) glitches to find correlation
- ✓ Fail to find long-duration (> few sec) glitches because those are always in coincidence with something happening in the witness channels

Deep-learning algorithms: in principle able to learn and evidence non-linear couplings