MACHINE LEARNING IN KM3NET

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THIRD ASTERICS – OBELICS WORKSHOP

New paths in data analysis and open data provision in Astronomy and Astroparticle Physics 23- 26 OCTOBER 2018

ASTERICS is a project supported by the European Commission Framework Progr Horizon 2020 Research and Innovation action under grant agreement n. 653477 TDOC CENTRE @ EDDINGTON 105 EDDINGTON PLACE CAMBRIDGE CB3 1AS

KM3Ne1

KM3NeT Infrastructure



Looking at neutrinos from below, ie. look towards the centre of the Milky Way Detector deployed deep under water

KM3NeT Experiment

KM3NeT-ARCA

Astroparticle Research with Cosmics in the Abyss

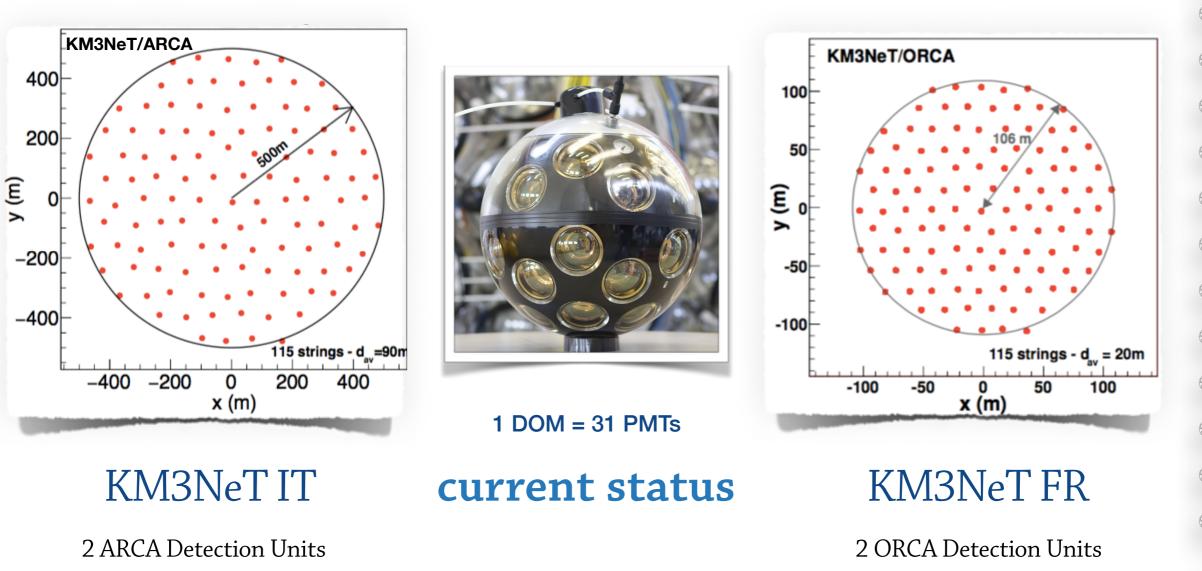
2 "building blocks" of 115 Detection Units
700m in height.
DOMs spaced 90m in X-Y, 36m in Z

KM3NeT-ORCA

Oscillation Research with Cosmics in the Abyss

Low energy neutrino detection - Neutrino mass hierarchy

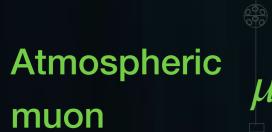
1 "building block" of **115** Detection Units: **200** m in height.
DOMs spaced **20**m in X-Y, **9**m in Z



 km^{3}

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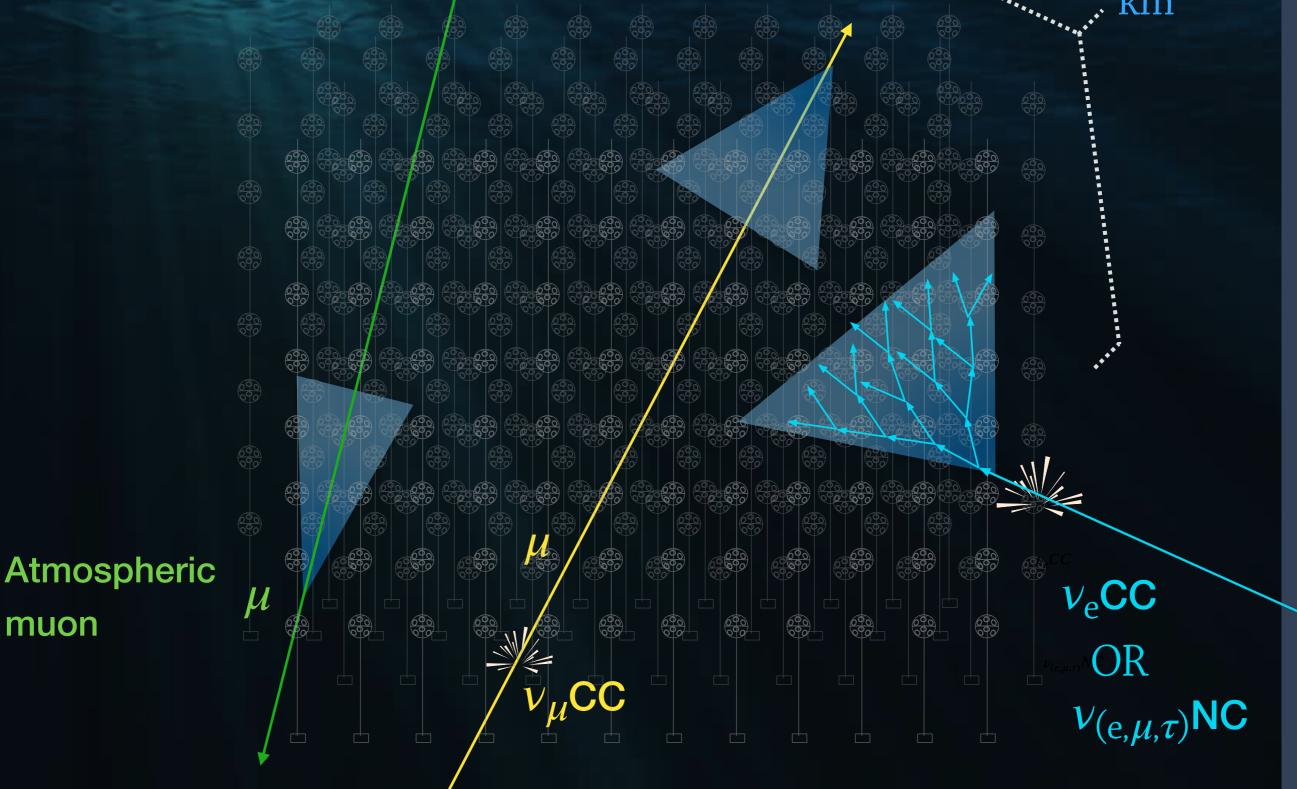
 km^3



 km^3

uCC

Atmospheric muon



 km^3

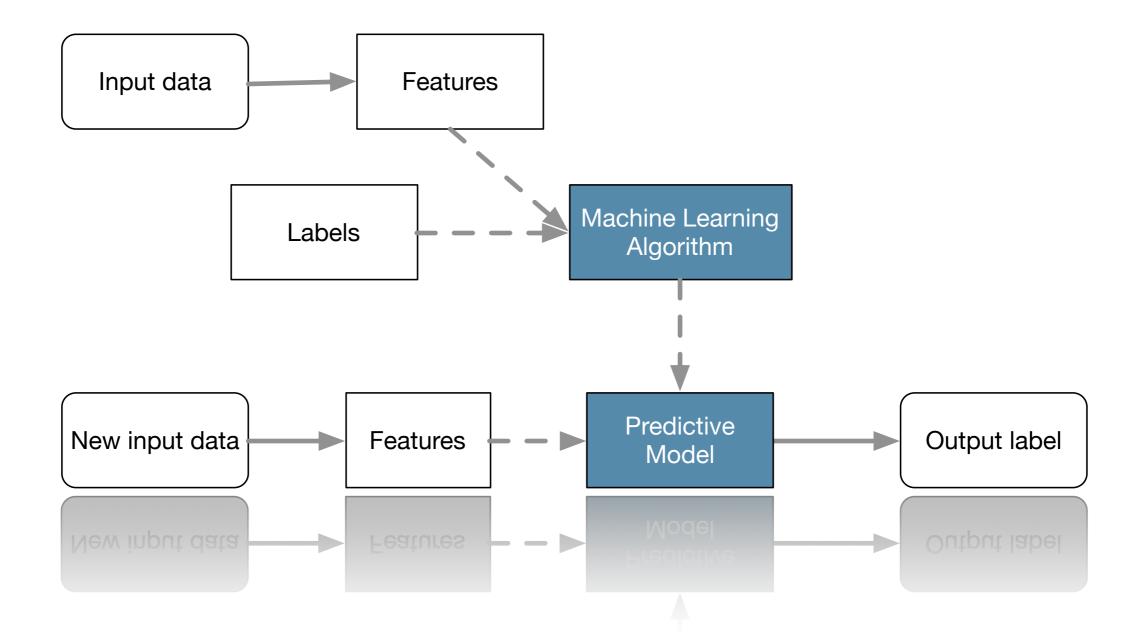
C. De Sio for the KM3Ne7 Collaboration – Third A OBELICS Workshop Cambridge

muon

STUDYING NEUTRINO INTERACTIONS WITH DEEP LEARNING

EXTRACT INFORMATION FROM RAW DATA - NO RECONSTRUCTION REQUIRED

Machine Learning: Learning from examples



Categorical labels: Classification

Continuous labels: **Regression**

(Some) Existing Machine Learning Applications for KM3NeT

- MVA in Point-Source Analysis (Credits A. Trovato)
 - · Random Forest for 3-class Prediction (Source, Atm $\nu,$ Atm $\mu)$
- High Energy Starting Muons (Credits K. Pikounis)
 - Boosted Decision Trees for 2-class Prediction (Signal, Background)
- EReNN: Energy Reconstruction with Neural Networks (Credits E. Drakopoulou, et al.)
 - Multi-Layer Perceptron for Energy Reconstruction
- Shallow and Deep Learning Applications in KM3NeT (Credits S. Geißelsöder et al.)
 - Multiple applications of Machine and Deep Learning Models in Supervised and Unsupervised Learning settings

Deep Learning applications for ARCA (C. De Sio - me) Deep Learning in ORCA with OrcaNet (M. Moser et al. - ECAP)

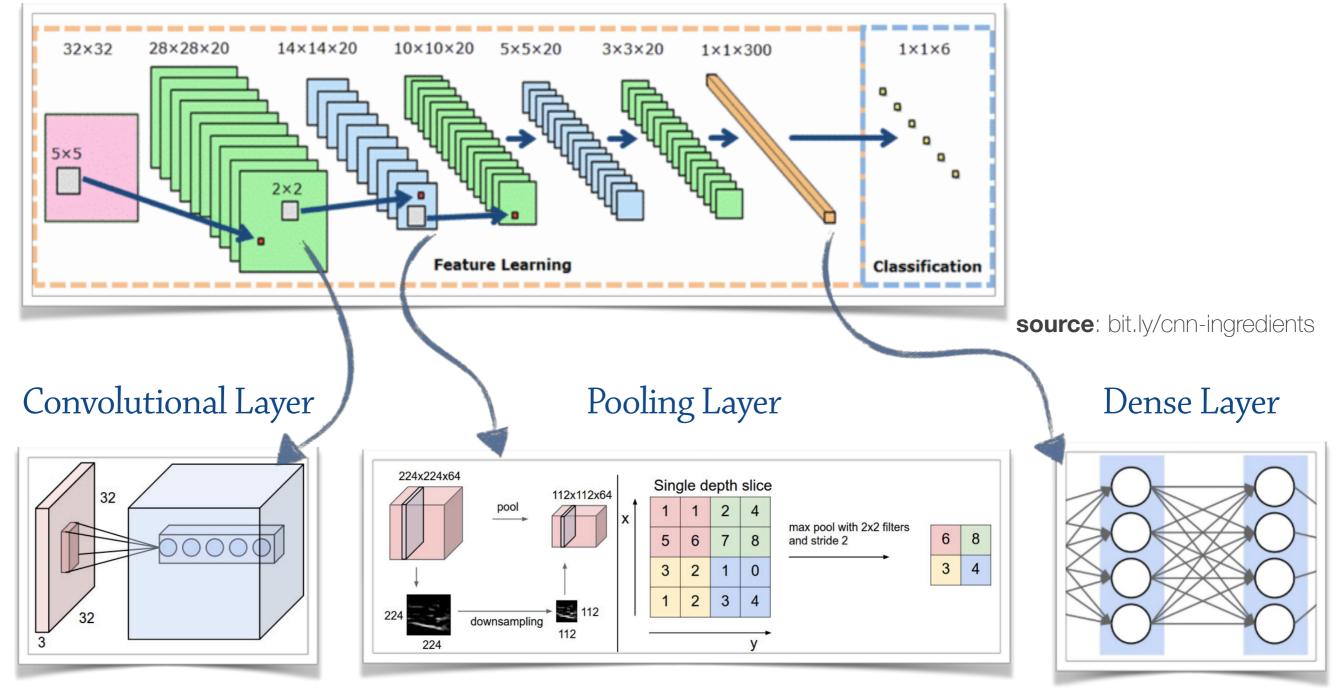
Deep Learning Applications for KM3NeT-ARCA

Four learning tasks:

- 1) Up-going/Down-going particle Classification
- 2) $v_{\mu}CC / v_eCC$ interaction Classification
- 3) Particle Energy Estimation
- 4) Particle Direction Estimation (Z component)

- Using **triggered hits** and times as input data
- Convolutional Neural Network models have been **designed** for each task

Convolutional Deep Networks (CNN): Main Ingredients

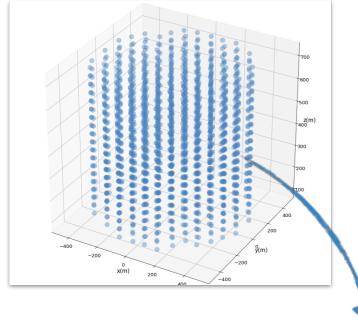


Local Feature Maps Learning Downscaling and Space Invariant Features Learning

Global Feature Learning & Prediction

DATA PREPARATION PREPARE DATA TO BE FED INTO NEURAL NETWORKS

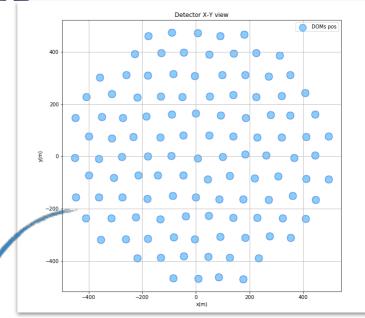
Space regularisation

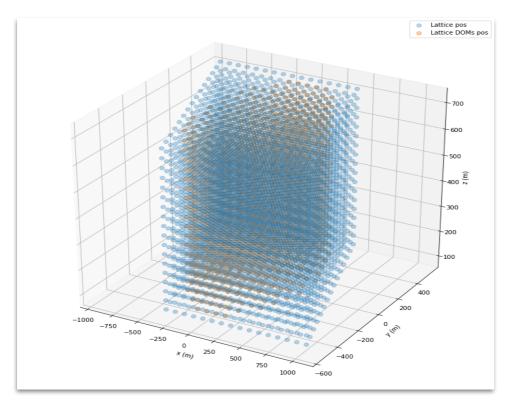


Regularised Detector Structure^{*}

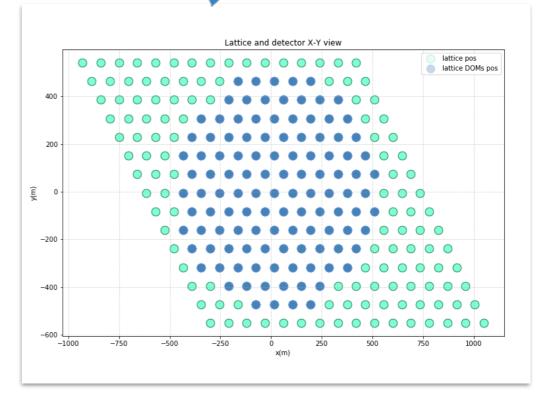
- **exactly** 90m spaced in (X,Y)
- **exactly** 36m spaced in Z

Regularised detector contained in **Lattice**





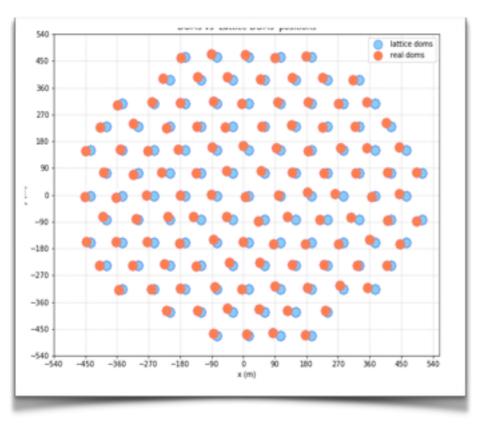




Regularised Detector XY-view

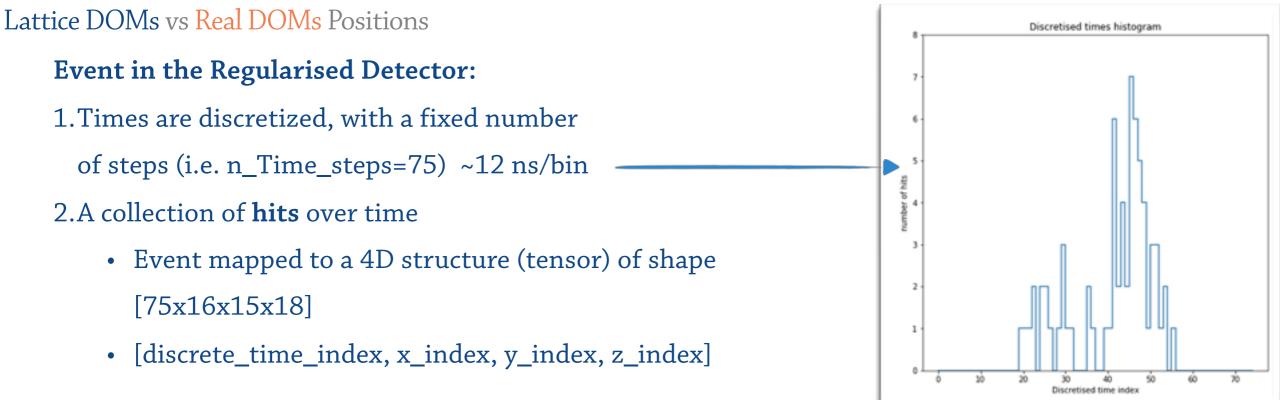
*Deviation from regularised structure can be introduced later as a next-order correction

Event Definition



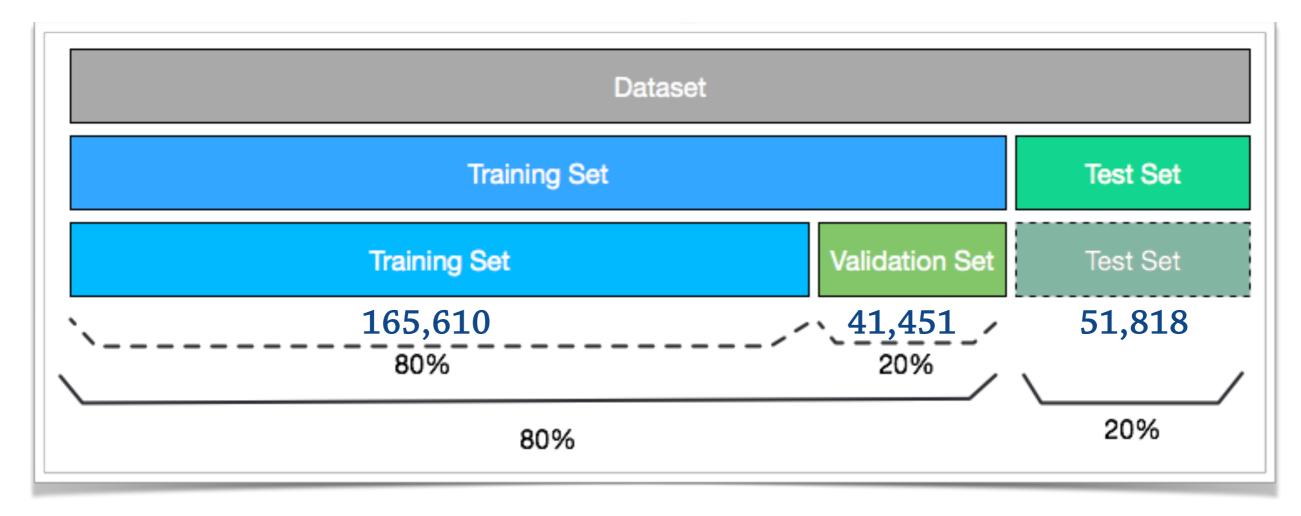
Reducing (useless and time consuming) sparsity in data

- Transforming DOM IDs into Lattice DOM IDs
 - a single DOM ID is mapped to an index in the [16x15x18] Lattice



Dataset

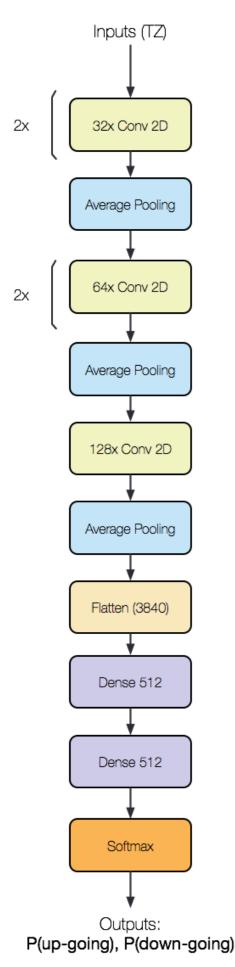
258,879 total events (samples) arranged from 100 v_eCC + 100 $v_{\mu}CC$ files





LEARNING TASK 1. UP-GOING/DOWN-GOING NEUTRINO CLASSIFICATION

CLASSIFY UP-GOING AND DOWN-GOING NEUTRINOS ACCORDING TO THEIR Z-COORDINATE EVOLUTION OVER TIME



Model architecture and Input Data

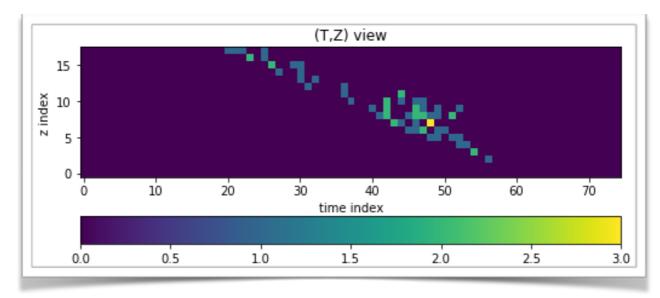
Labels

y: $\begin{cases} \cos(\theta_z) > 0 : "up-going" \\ \cos(\theta_z) \le 0 : "down-going" \end{cases}$

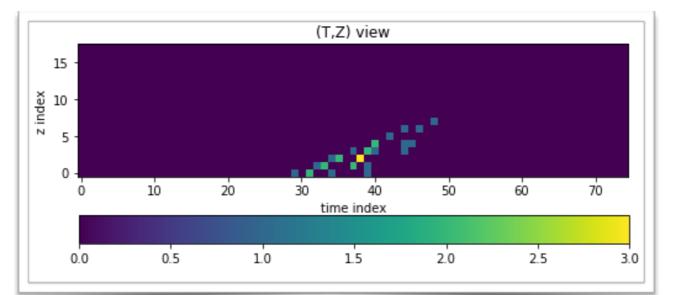
Input tensors reshaped to:

(T,Z): [n_samples,75,18]

Input array summed over X and Y axes



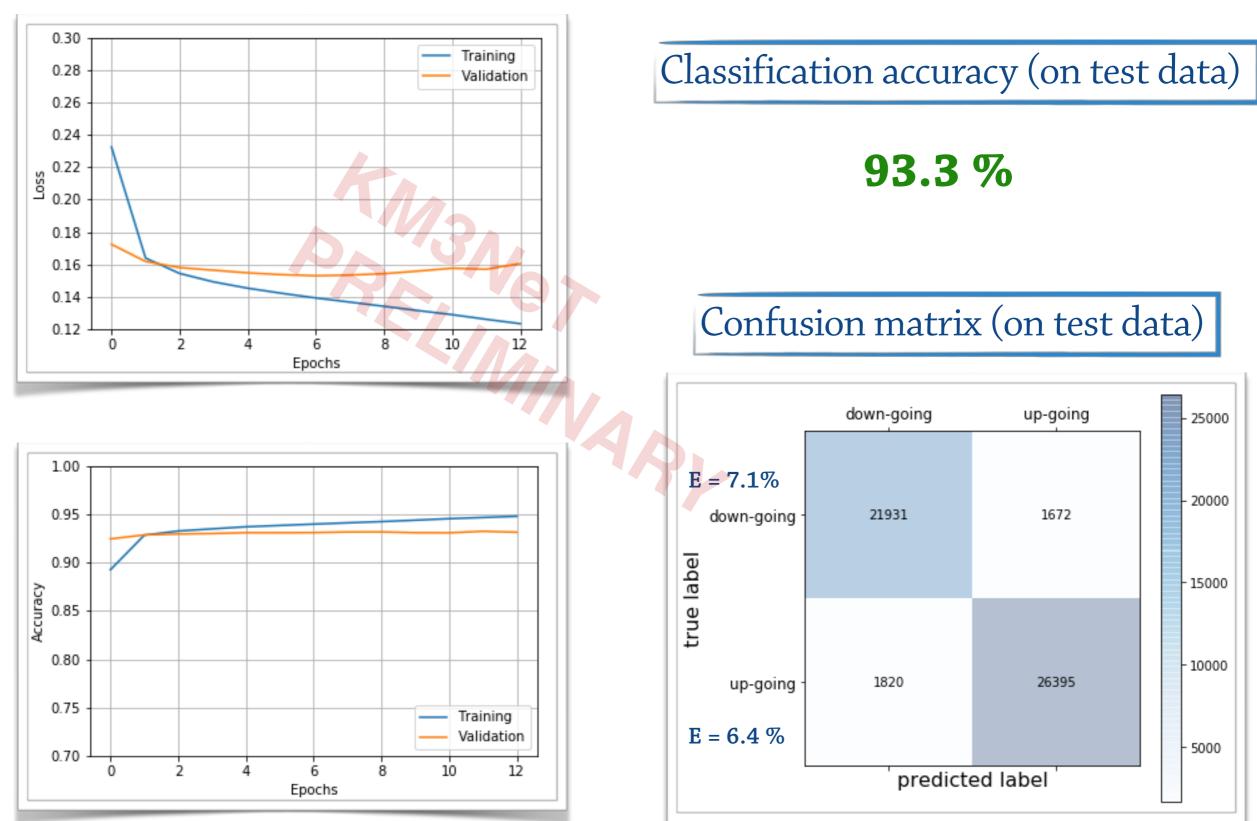




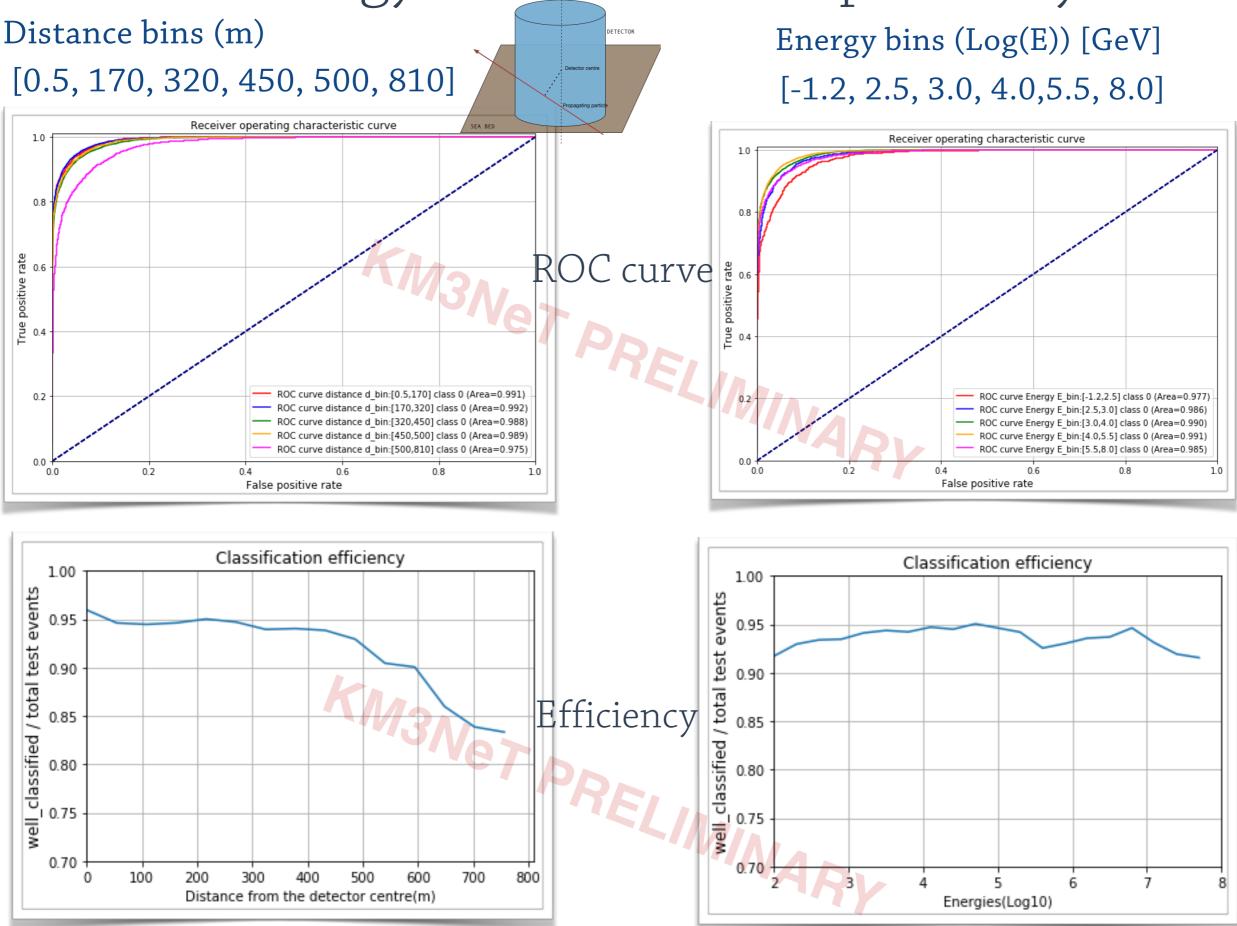
Up-going

Classification results

Training history



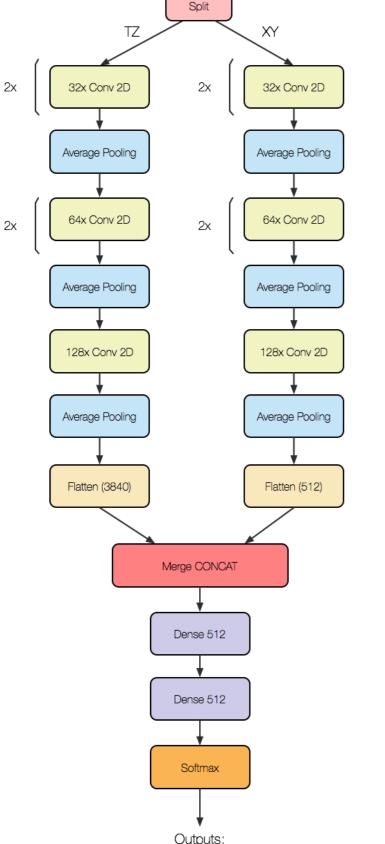




LEARNING TASK 2. $\nu_{\mu} CC / \nu_{e} CC$ **INTERACTION CLASSIFICATION**

CLASSIFY $\nu_{\mu}CC / \nu_{e}CC$ **INTERACTIONS BASED ON** THE SHAPE OF THE EVENTS AND THE EVOLUTION OF POSITIONS OVER TIME

Model architecture and Input Data



Outputs: $P(v_{\mu}), P(v_{e})$

CNN model with parallel branches analysing (T,Z) and (X,Y) evolution separately, merged to extract common features

2 Input tensors of shapes:

(T,Z): [n_samples, discrete_time_index, z_index]

(X,Y): [n_samples, x_index, y_index]

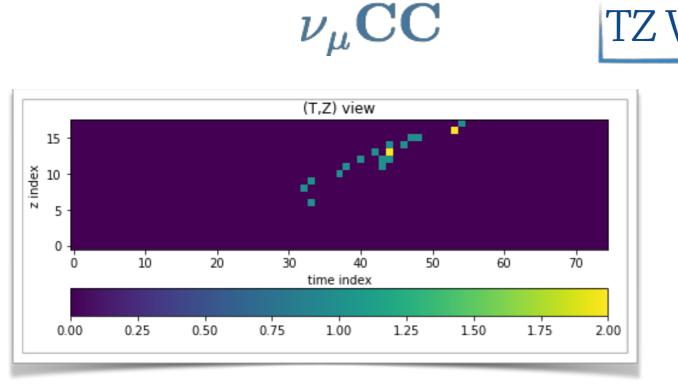
Labels

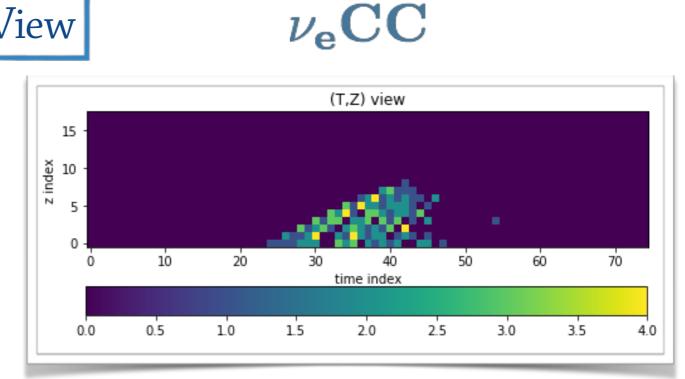
$$= \begin{cases} 0: v_{\mu}CC \\ 1: v_{e}CC \end{cases}$$

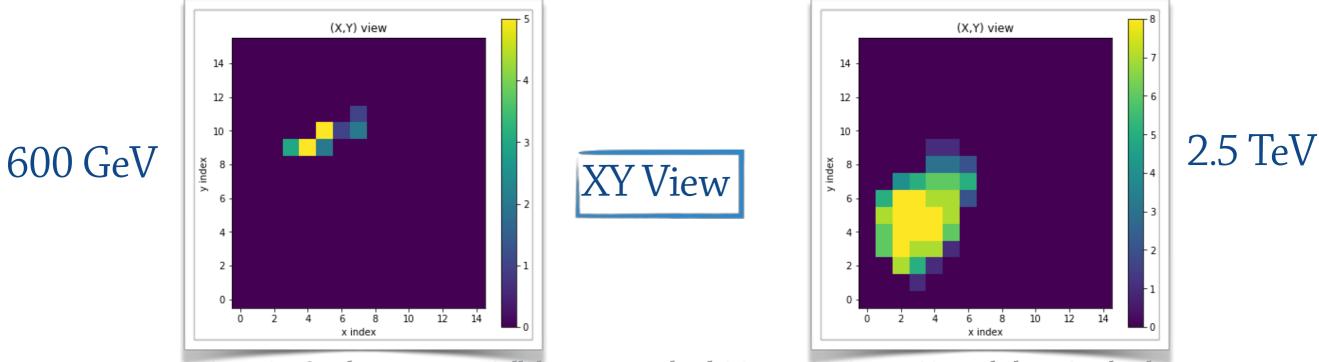
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Multiple Views of Events

TZ View

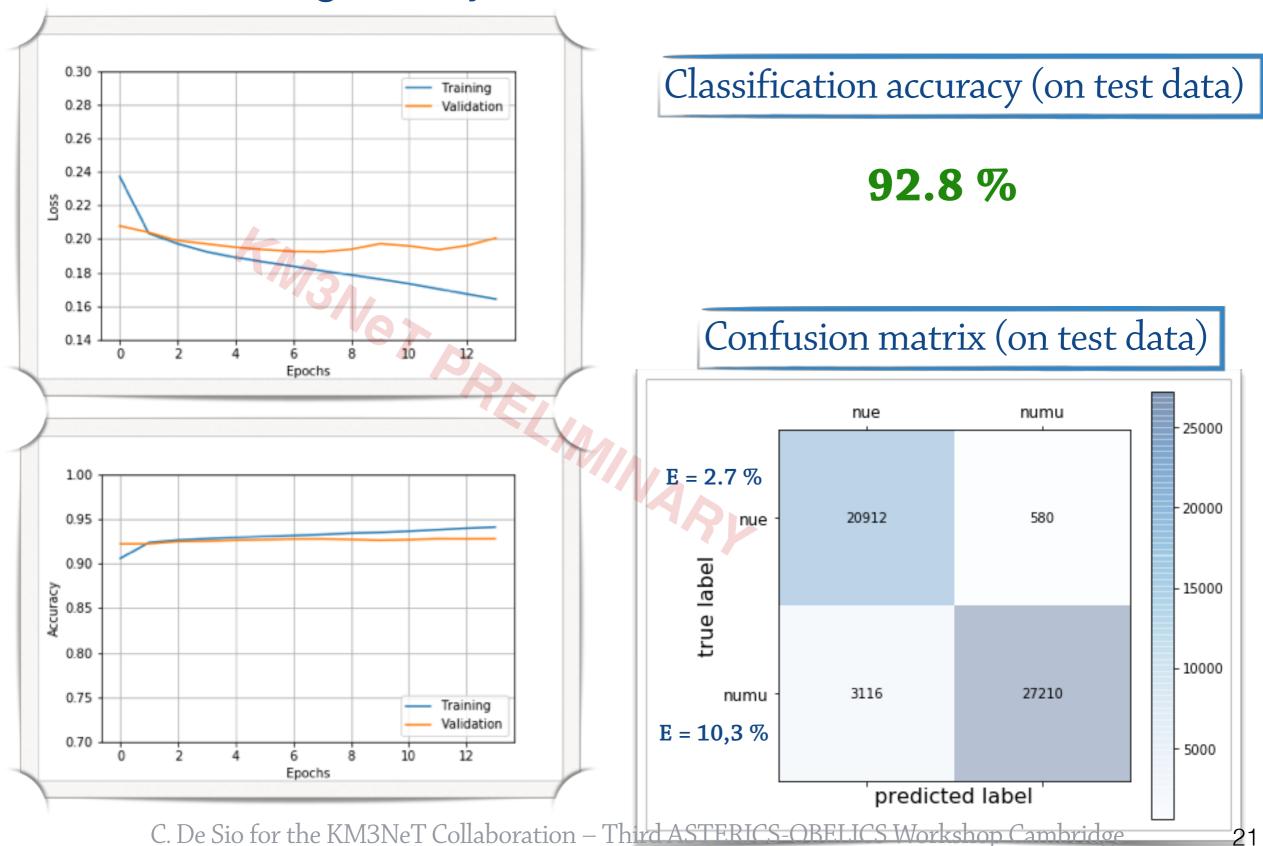






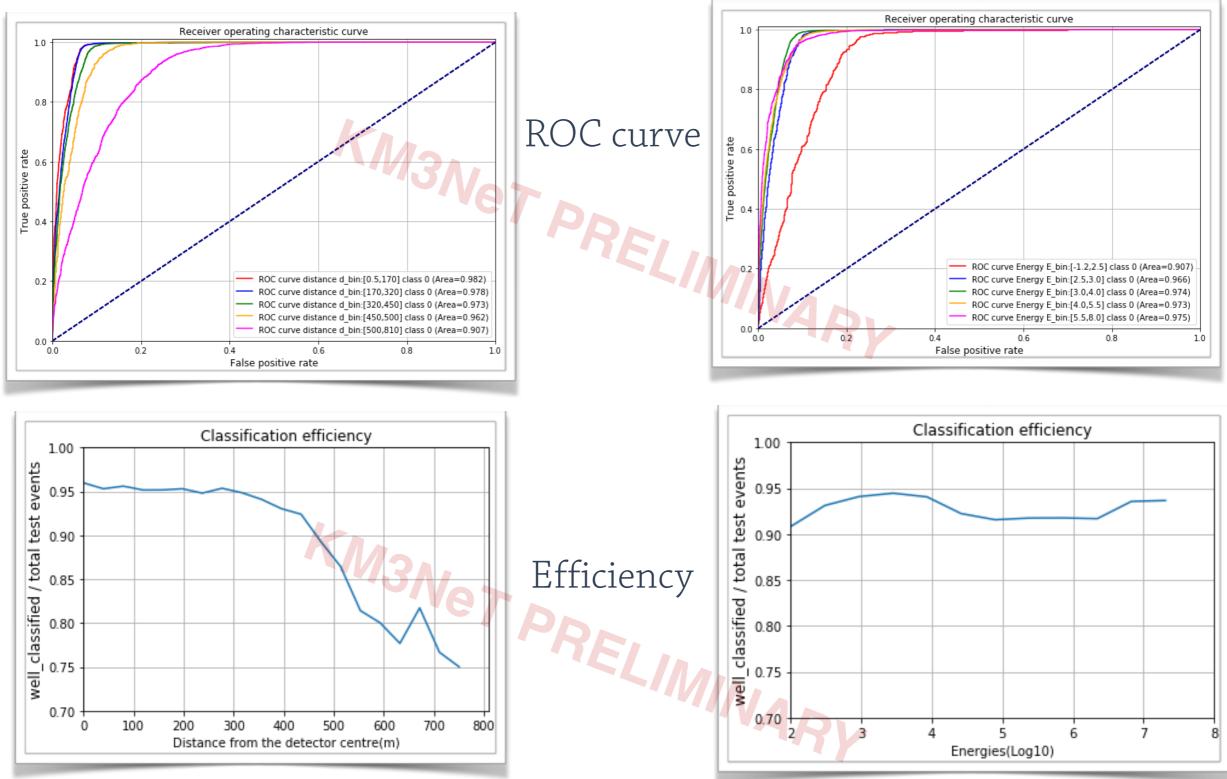
Classification results

Training history



Energy and Distance dependency

Distance bins (m) [0.5, 170, 320, 450, 500, 810] Energy bins (Log(E)) [GeV] [-1.2, 2.5, 3.0, 4.0, 5.5, 8.0]



LEARNING TASK 3. PARTICLE ENERGY ESTIMATION REGRESSION MODEL TO ESTIMATE NEUTRINO ENERGY (GEV)

Model architecture and Input Data Inputs (TXYZ)

CNN model with parallel branches analysing (T,Z) and (X,Y) evolution separately, merged to extract common features, fed into multiple fully connected layers

2 Input tensors of shapes:

(T,Z): [n_samples, discrete_time_index, z_index]

(X,Y): [n_samples, x_index, y_index]

Labels:

Split

Merge CONCAT

Dense 128

Dense 64

Dense 32

Dense 16

Linear

Outputs: LogE

2x

2x

32x Conv 2D

Average Pooling

64x Conv 2D

Average Pooling

128x Conv 2D

Average Pooling

GlobalAvaPoolina (128)

XY

32x Conv 2D

64x Conv 2D

128x Conv 2D

Average Pooling

BlobalAvgPooling

(128)

Energy: MC truth

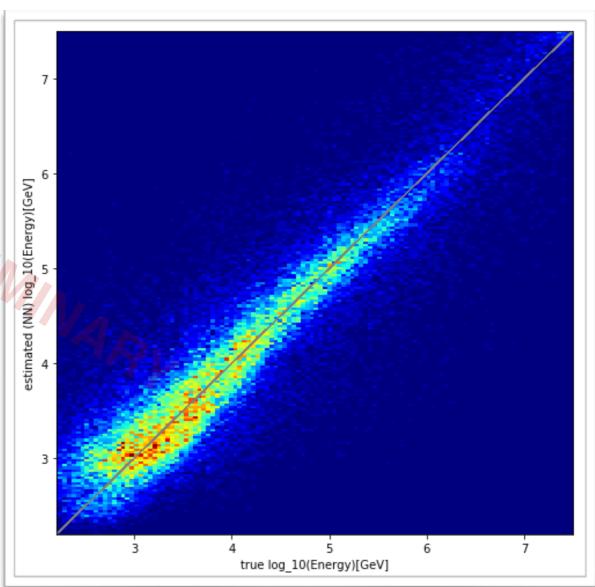
Results

Training history 0.40 Training Validation 0.35 7 0.30 Loss estimated (NN) log_10(Energy)[GeV] 0.25 0.20 20 15 10 0 5 Epochs Mean Squared Error (on test data) 3 0.22

r² score (on test data)

0.84

Estimated vs True Energy

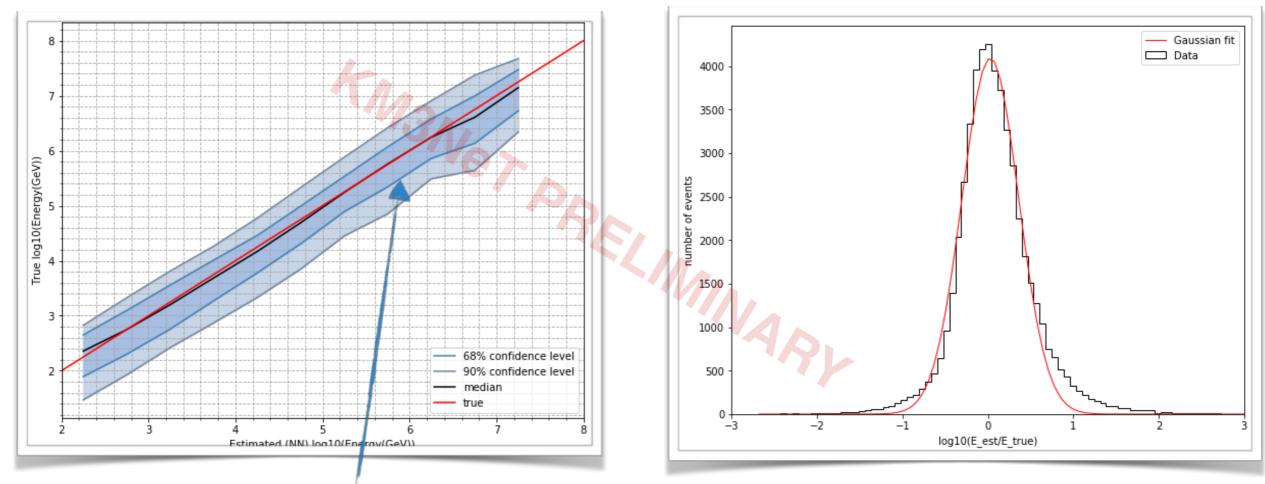


test dataset ($\nu_{\mu}CC + \nu_{e}CC$ events)

Results

True vs Neural Network Estimated Energy

Energy resolution



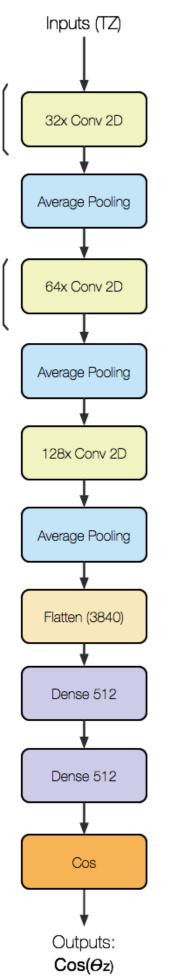
lack of statistics at high energy

Gaussian fit w/ μ = 0.03, σ = 0.33

test dataset ($\nu_{\mu}CC + \nu_{e}CC$ events)

LEARNING TASK 4. DIRECTION ESTIMATION (Z)

REGRESSION MODEL TO ESTIMATE Z-COMPONENT OF THE NEUTRINO DIRECTION



Model architecture and Input Data

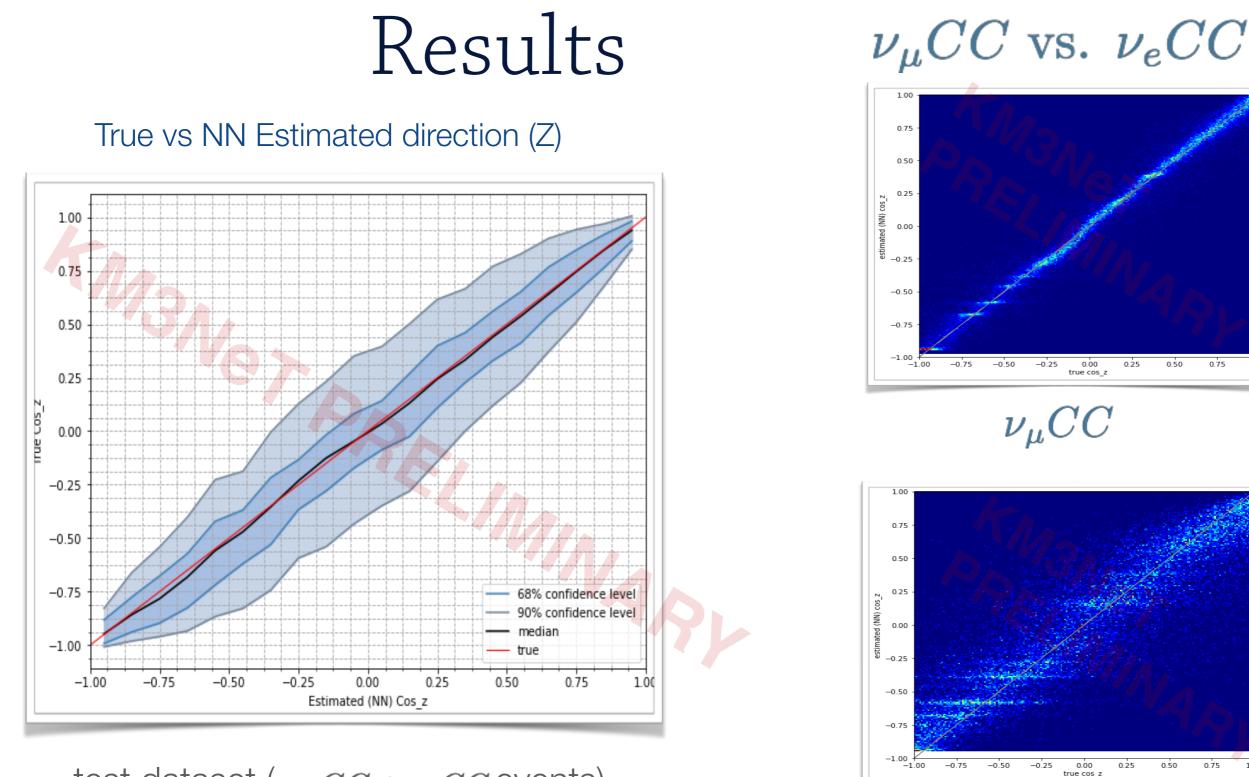
CNN model to analyse (T,Z) evolution and predict $cos(\theta_z)$ value

Input tensor reshaped to :

(T,Z): [n_samples, discrete_time_index, z_index]

Labels

 $\cos(\theta_z)$: MC truth



test dataset ($\nu_{\mu}CC + \nu_{e}CC$ events)

Mean Squared Error (on test data)

0.03

Muon neutrino event direction is better estimated w.r.t. electron

 $\nu_e CC$

-0.50

-0.25

0.25

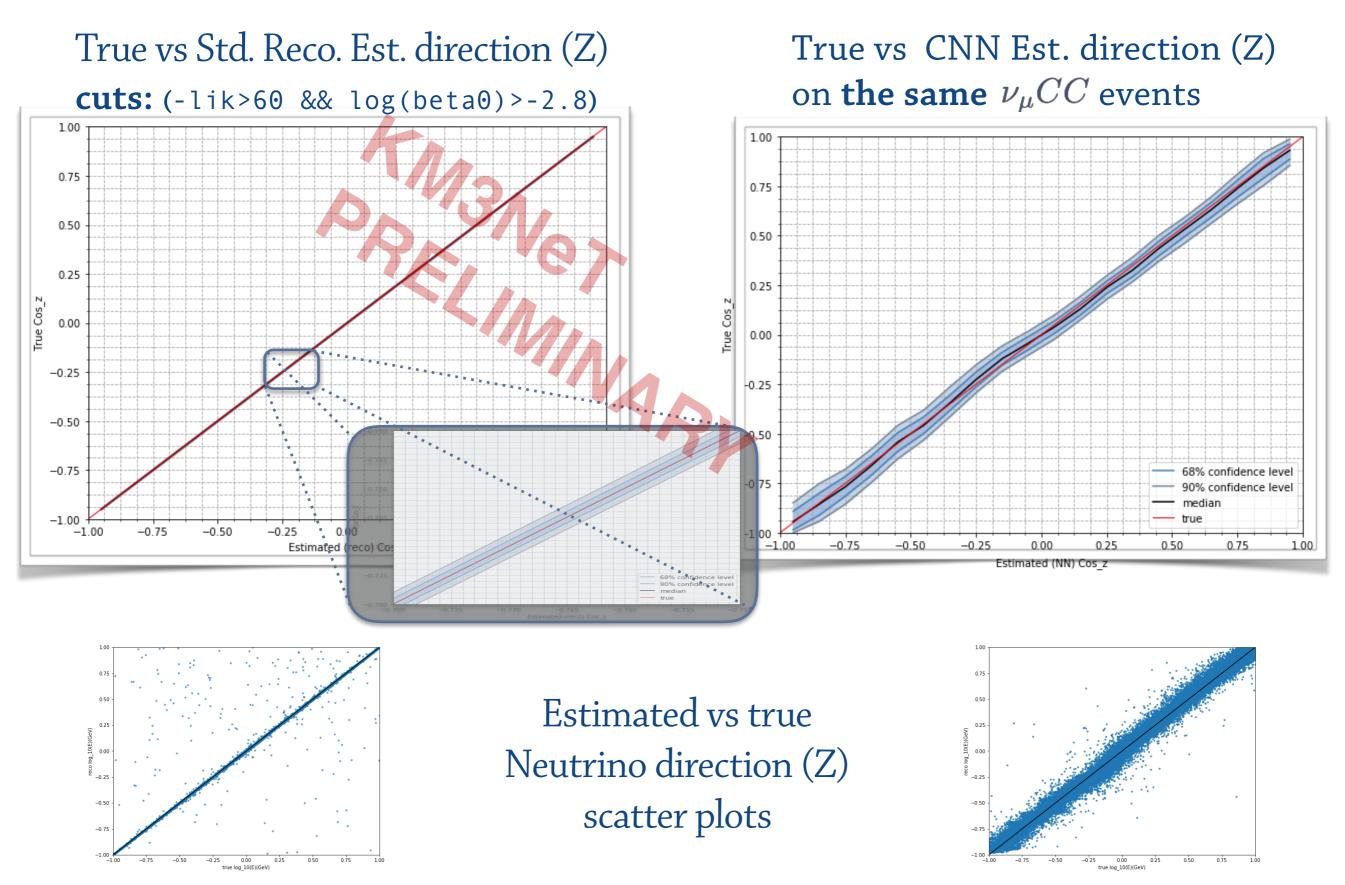
0 50

0.75

COMPARISON WITH THE KM3NET STANDARD RECONSTRUCTION ALGORITHM

PERFORMANCES COMPARISON ON $\nu_{\mu}CC$ EVENTS TO COMPARE RESULTS WITH STANDARD TRACK RECONSTRUCTION ALGORITHM, WHEN APPLICABLE

Direction estimation (Z)



Comparison on Up-going/Down-going classification

- $\cdot \;$ apply "labels" to reconstructed events:
 - · $\cos(\theta_z) > 0$: "up-going"
 - $\cos(\theta_z) <= 0$: "down-going"
- Compare predictions

Accuracy on up-going/down-going classification

KM3NeT standard reconstruction (-lik>60 && log(beta0)>-2.8)

CNN regression running test on $v_{\mu}CC$ events only

Classification Accuracy

0.998

Classification Accuracy

0.987

only $v_{\mu}CC$ events (selecting well reconstructed events) with quality cuts

Deep Learning Applications for KM3NeT-ORCA

Aim:

Multipurpose classification and regression studies in ORCA

Task:

Reconstruct energy and direction, track-shower composition, PID

 Developed a Deep Learning track-shower classifier called OrcaNet, based on Convolutional Neural Networks

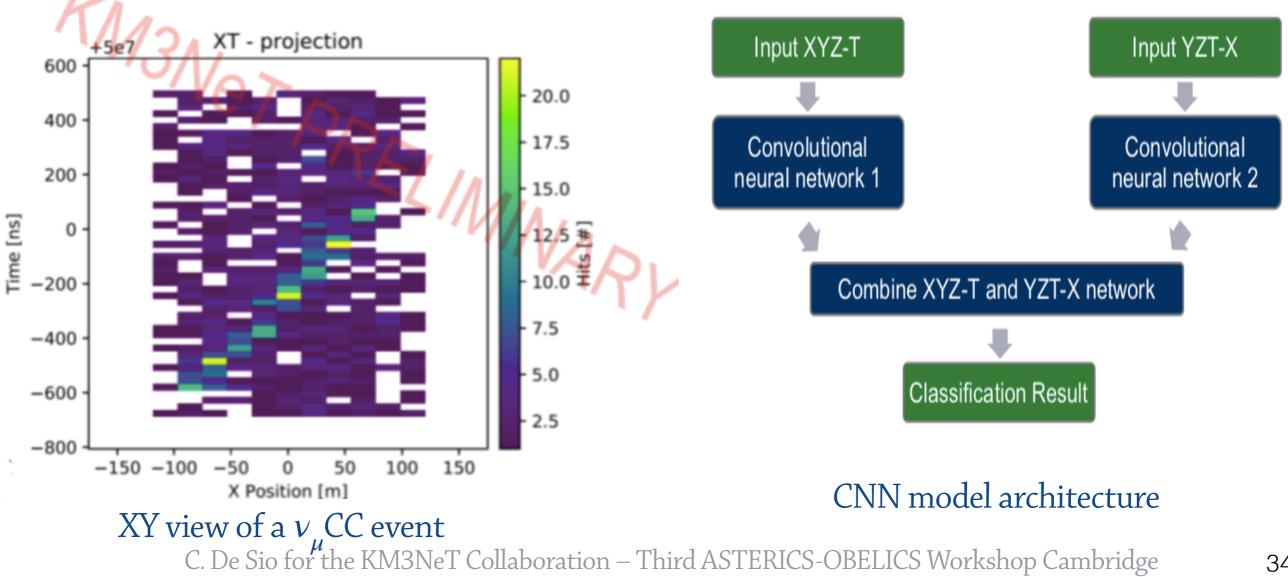
Deep Learning for ORCA with OrcaNet

6D information (XYZ, T, azimuth, zenith) projected to 4D subspace (neglect PMT direction)

Time binning: 60 bins (~10 ns/bin)

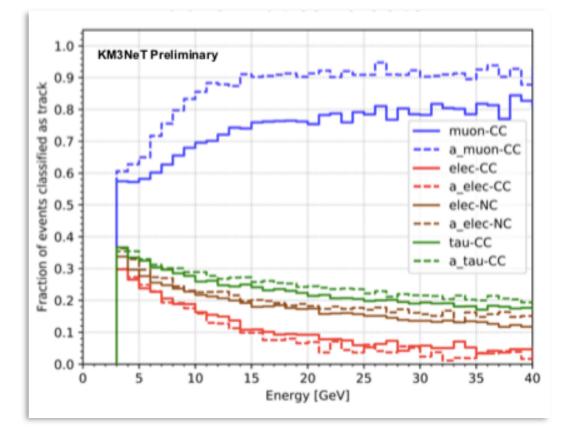
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- Space binning: 11x13x18
- 4 week training on HPC GPU Cluster (4xGTX 1080)

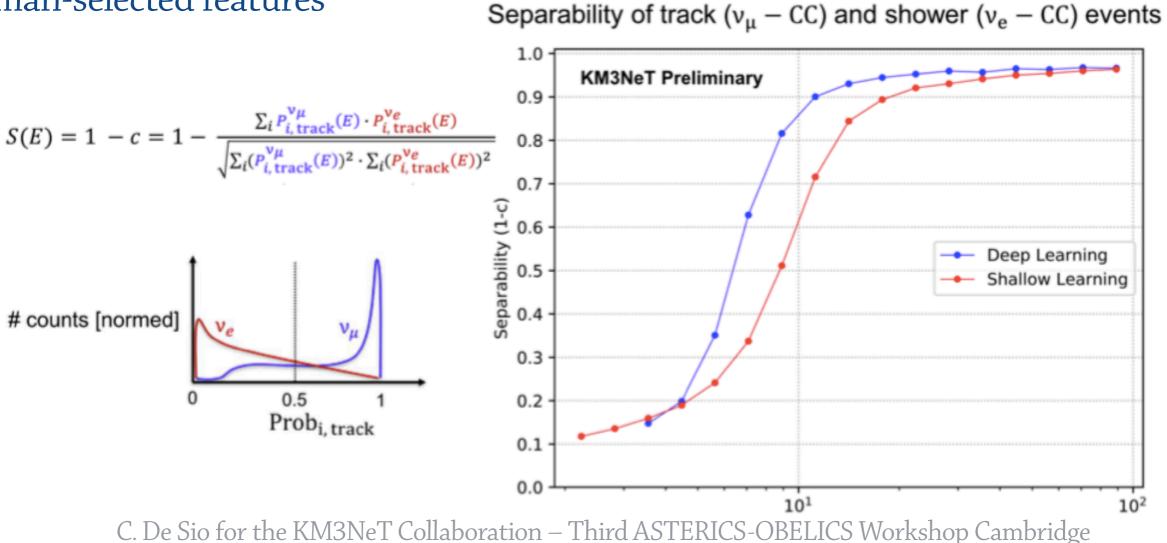


Results

Track/shower event classification No reconstruction output used



CNN **outperforms** RDF with human-selected features



35

Conclusions and outlook



- ML Results are promising in these **first iterations**:
 - ML provides stable estimations, comparable and in some cases better than KM3NeT standard reconstruction
 - ML does not depend on any reconstruction algorithm: independent event study directly from raw data
- Room for improvement and a lot of work to do (detailed detector description, complete reconstruction, complete flavour identification) – More results coming soon!
- KM3NeT expects to produce a DL toolset for CNN applications in ASTERICS (H2020) WP 3
 Possibly portable to other event-based experiments, to be added to ASTERICS repository
- KM3NeT INFN groups to contribute with ML algorithms in Task 3.4 of ESCAPE (proposal)

Thanks a lot for your kind attention

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