Radio-optical synergies: Extracting multi-wavelength science

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Focus of the data school so far



Radio spectrum is ~ a power law

Redshift completely unconstrained

Radio data alone offers limited scientific value

Need multi-wavelength information to maximise the scientific potential



Motivation and Overview

Radio alone offers limited scientific use -> need to extract multi-wavelength properties

- Host-galaxy (or multi-wavelength counterpart) identification
- (Photometric) redshift estimation
- A demonstration/example use case on a LOFAR dataset

* Colloquially: "opical" = wavelengths other than radio

* Caveat: A focus on galactic/extragalactic science, but applicable to other cases

Radio-optical cross-match

The LOFAR surveys

LOFAR Two-meter Sky Survey (LoTSS) DR2



Data release 2: > 4 million radio sources! Covering > 5500 deg²

LoTSS DR2 - Multi-wavelength datasets

DESI Legacy Imaging Surveys:

Optical g, r, z photometry

+unWISE IR deblended photometry



LoTSS Deep Fields (DR1)

- Deepest wide-field radio continuum survey to date at low frequencies
- 100s hrs of radio imaging over 25 sq. deg.
- 3 fields: ELAIS-N1, Lockman Hole, Bootes

Tasse et al. 2021; Sabater et al. 2021 Kondapally et al. 2021; Duncan et al. 2021; Best et al. 2023



LoTSS Deep Fields

Dec

ELAIS-N1

Excellent deep, wide-area coverage from UV to far-IR (~2 million sources)

+ radio (~30,000 sources)

~ 7 deg² of overlap



Cross-matching: a challenging endeavour

Fairly straight-forward *if sources are compact/unresolved*





However...

Cross-matching: a challenging endeavour

Radio source population is diverse, with complex emission and morphologies -> difficult to simply use NN methods

Source detection software don't know astrophysics!

 Complex emission -> physical sources not correctly identified -> requires manual intervention

Radio image with PyBDSF sources



Cross-matching: a challenging endeavour

Source blending is also an issue (esp. for deep surveys)

Image: three distinct physical sources, categorised as a *single* PyBDSF source

Radio contours + IR image



Williams et al. 2019; Hardcastle et al. 2023

Combination of statistical (automated) method for simple sources and visual classification for complex sources

Statistical Likelihood Ratio method

Probability of a galaxy to be the genuine counterpart to given radio source

$$LR = \frac{f(r) \ q(m,c)}{n(m,c)}$$

Visual Classification

Visual classification used for more complex sources

Both host-galaxy identification and source association

Using the Zooniverse framework

Consensus decision from at least 5 astronomers

* Different process for LoTSS DR2

Radio contours + IR image



Williams et al. 2019; Hardcastle et al. 2023

When to use LR versus visual classification?

Use a **decision tree** to identify when to use LR versus visual classification

In LOFAR Deep Fields:

Counterparts found for > 97% of radio sources

LR: 85% of radio sources

Visual classification: 15%



Kondapally et al. 2021

Science applications utilising multi-wavelength information

Spectroscopy provides most robust measure of redshifts and many other galaxy properties

Photometric Redshifts - two ways

Template Fitting

EAZY Brammer et al. (2008) LePhare Arnouts et al. (1999) Ibert et al. (2006)

PhotoZ Bender et al. (2001) **Hyper-Z** Bolzonella et al. (2000)

ZEBRA Feldmann et al. (2006) **Mizuki** Tanaka et al. (2015)

BPZ Benitez (2000)



Photometric Redshifts - two ways

Machine Learning

Generalized Linear Models Radial basis function networks Self-organizing Maps (SOMs) **Bayesian Network** Naive Bayes Gaussian Processes **Neural Networks** Deep learning **Randomised Forests** k-Nearest Neighbour **Boosted Decision Trees** Relevance vector machines Support Vector Machines (SVM) Normalised inner product nearest neighbour Directional neighbourhood fitting Non-conditional density Voronoi tesselation density estimator estimation

Diverse radio source population

Diverse nature of the radio source population

-> Single approach may not be most suitable



Best et al. 2023

Photometric Redshifts: a hybrid approach

A hybrid approach within LOFAR surveys for photo-z's – led by Ken Duncan



Duncan et al. 2018a,b; Duncan et al. 2021; Duncan 2023 Spectroscopic z (gold standard; "ground truth")

The future: new spectroscopic redshift surveys



WHT 4.2m Telescope 370-960nm wavelength coverage 960 fibres @ R~5000 Northern hemisphere *Starting Q2 2024*



VISTA 4m Telescope 370-960nm wavelength coverage ~1400 fibres @ R~5000 Southern hemisphere *Starting 2025*



Mayall 4m Telescope 360-980nm wavelength coverage 5000 fibres @ R~2-5000 Northern hemisphere *Started 2022-23*



Subaru Telescope 0.38-1.2micron coverage 2400 fibres @ R~2-5000 Northern hemisphere *Starting >2024*



VLT 8m Telescope 0.68-1.8micron coverage ~1000 fibres @ R~4-6000 Southern hemisphere *Starting ~Q2 2025*



WEAVE-LOFAR (1.3m fibre hours) WEAVE-Apertif (Resolved spectroscopy of HI sources)



WAVES (GAMA Mrk.2) Cosmology Redshift Survey AGN Survey (eROSITA sources) Cluster survey ORCHIDDS (Optical Radio and HI Deep Spectroscopic Survey; PI Duncan)



Cosmology survey: 'Bright galaxies' at z < 0.4 Luminous Red Galaxies Emission line galaxies (ELGs) Quasars



Cosmology: ELGs at 0.6 < z < 2.4 Galaxies: near-IR selected deep fields



Deep near-IR selections (e.g. COSMOS/CDFS etc) Some MIGHTEE follow-up (TBC)

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WEAVE

WHT Enhanced Area Velocity Explorer

- WEAVE is a ~1000 fibre multi-object spectrograph going on the 4.2m WHT
- 2 deg diameter field of view
- Complete wavelength coverage from 370–960nm at R=5000
- Starting Q2 2024!





WEAVE-LOFAR

Spectroscopic follow-up of large numbers of LOFAR-selected sources in three tiers, to get a complete picture of SF and AGN co-evolution.



Demonstration of combining radio and multi-wavelength datasets

A giant radio galaxy example



Legacy Survey Viewer + Catalogue

https://www.legacysurvey.org/viewer



Legacy Survey Viewer + Catalogue

https://www.legacysurvey.org/viewer



PanSTARRS Image Cutout Service

https://ps1images.stsci.edu/cgi-bin/ps1cutouts



PanSTARRS-1 Image Access

 170.3577167 53.7492520
 Submit
 Reset
 Clear
 Help

 Filters:
 Image: Clear
 Image: Clear
 Image: Clear
 Help

 Filters:
 Image: Clear
 Image: Clear
 Image: Clear
 Help

 Filters:
 Image: Clear
 Image: Clear
 Help

 Auxiliary data:
 Image: Clear
 Help

 Auxiliary data:
 Image: Clear
 Help

 Cutout image size:
 5000
 pixels (1250.00 arcsec) (sets spatial size of the FITS image)

 JPEG display size:
 256 v pixels (sets resolution of the JPEG previews)

170.3577167 53.7492520 (ra = 170.357717, dec = 53.749252)



MAST/STScl

LOFAR 150 MHz

PS1 (g-band)



Conclusions

- Need to incorporate multi-wavelength data to maximise scientific potential from radio surveys
- Methods of cross-matching radio and optical surveys
- Estimating photometric redshifts
- Useful resources for finding and obtaining multi-wavelength datasets for your radio data

Radio-optical analysis: useful resources



Radio-optical analysis: useful resources

PanSTARRS (PS1) DR1 and DR2

Optical g,r,i,z,y bands

Images and catalogues: https://outerspace.stsci.edu/display/PANSTARRS/

Legacy Surveys (DR10 latest)

g,r,z bands (deeper than PS1) + unWISE (IR) de-blended photometry Measurements of redshifts, simple galaxy profiles/sizes

Images and catalogues: <u>https://legacysurvey.org</u>

<u>SDSS</u>

Large optical spectroscopic data

Many data releases, (mostly) local Universe, spec-z's, galaxy properties https://www.sdss4.org/

Radio-optical analysis: useful resources

Access to large IR datasets NASA IR Science Archive (IRSA):

https://irsa.ipac.caltech.edu/

Interactive sky viewer with many overlays and easy data access functionality Legacy Sky Viewer: https://www.legacysurvey.org/viewer

Large public catalogue repository with flexible access, searching, cross-match Vizier/CDS: <u>https://vizier.cds.unistra.fr/viz-bin/VizieR</u>

Space telescopes archive **MAST:** <u>https://archive.stsci.edu/</u>

Interactive viewer for (most) public data ESA Sky: <u>https://sky.esa.int/esasky/</u>

> NASA Extragalactic Database (NED): https://ned.ipac.caltech.edu/

