

GPU-based correlation for the EVN

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Introduction

- Current EVN cluster: 840 Intel Xeon cores (2018)
 - 20 stations @ 2 Gpbs in realtime
 - Uses 7.5 kW of power!
- Next gen receivers, tri-band, C/X, ... : 4 GHz+ bandwidth
- Next generation backends, DBBC3, DBBC4,.. : 32 Gbps+
- Moore's law for CPUs is dead
- Next generation EVN correlator needs to green
- GPUs can provide an order of magnitude of improvement in efficiency
- Funding in RADIOBLOCKS to develop GPU powered VLBI correlator

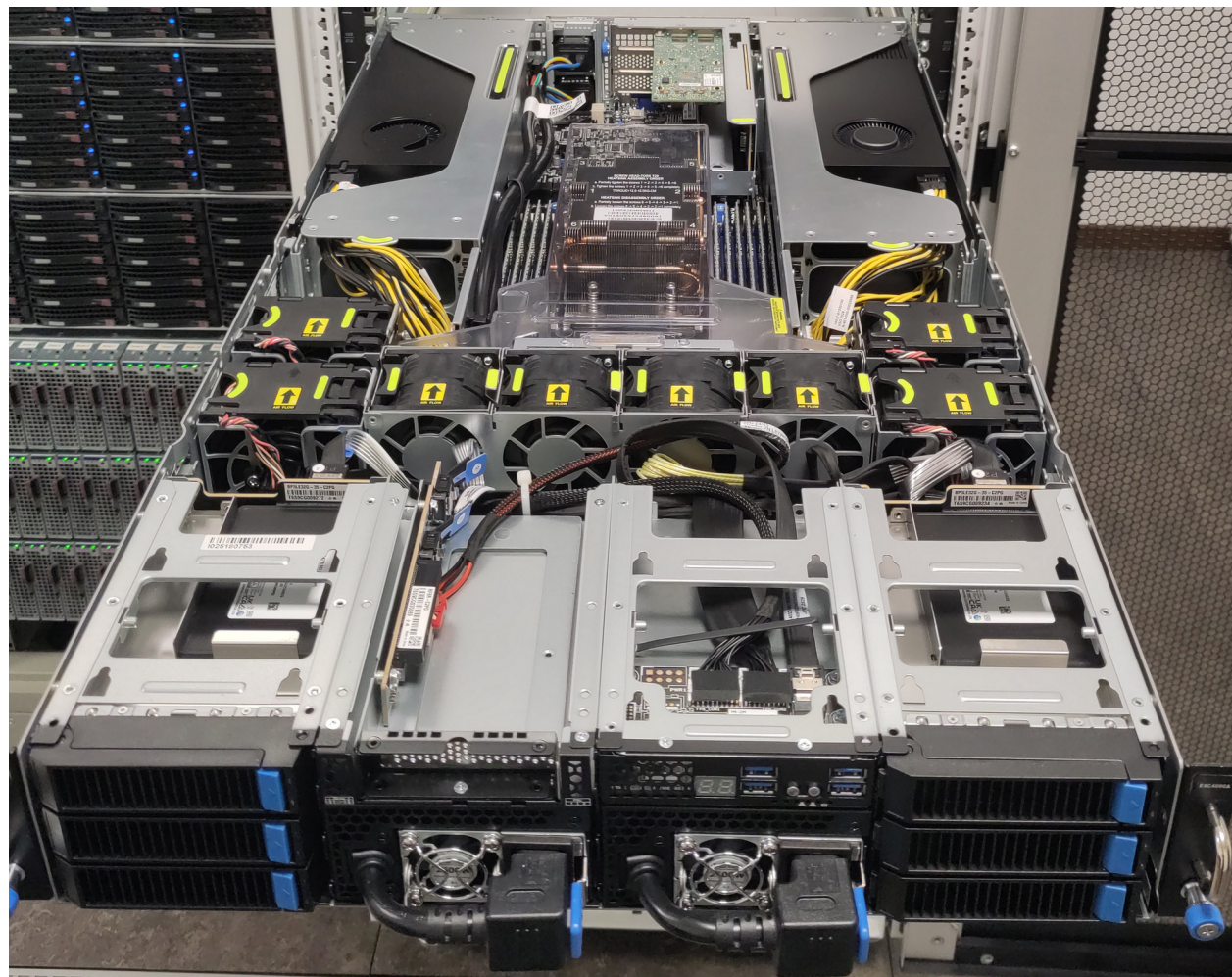
SFXGPU

- New correlator, not directly based on SFXC
- Same correlation algorithm as SFXC
- Can already correlate 95% of EVN experiments
- Implemented features include
 - Mixed-bandwidth correlation
 - Hanning smoothing
 - Multiple simultaneous phase centres
 - Coherent de-dispersion
- To-do
 - Pulsar folding
 - Geodetic VLBI support (phase-cal extraction)
 - eVLBI support
- Source code: **<https://code.jive.eu/keimpema/sfxgpu>**

GPU correlator

- **CPU:** 1x MD Epyc 9555P - 360W
- **GPU:** 4x **NVIDIA RTX PRO 4000** PNY
Blackwell 24GB GDDR7
PCI Express x16 5.0 - 140W
- **RAM:** 768 GB (12x 64GB 4800MT/s
DDR5 ECC Reg CL40 DIMM)
- **Storage:** 2x .9TB SSD U.2 NVMe
PCIe 4.0 x 4 Enterprise
- **NIC:** ConnectX-6 DX EN adapter card,
100GbE, Dual Port

*Funded through the Radioblocks
project*



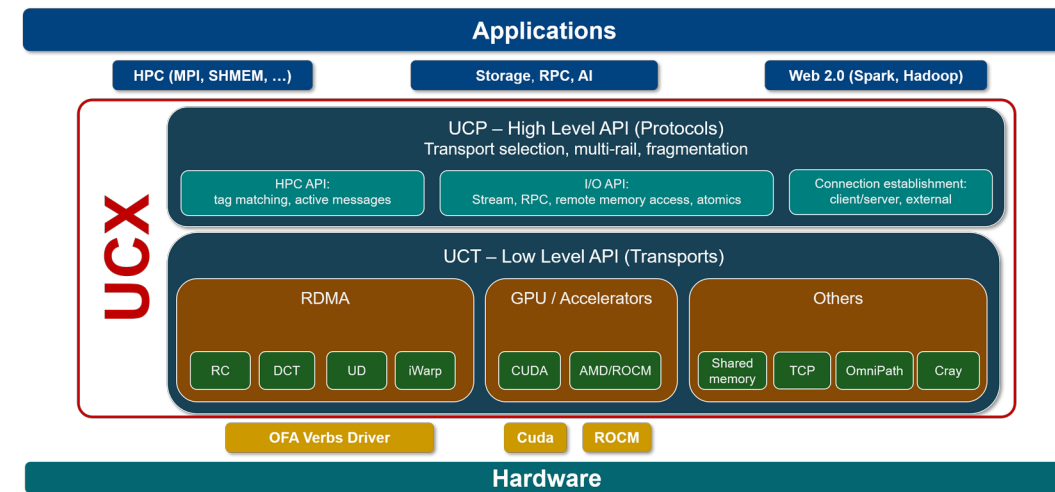
Bye Bye MPI; Welcome UCX

Abandoned MPI in favour of (lower-level) UCX library

- Easier to deploy in heterogeneous environment
- Use wakeup feature to save energy (no polling)

Split correlator application into separate reader and correlator processes

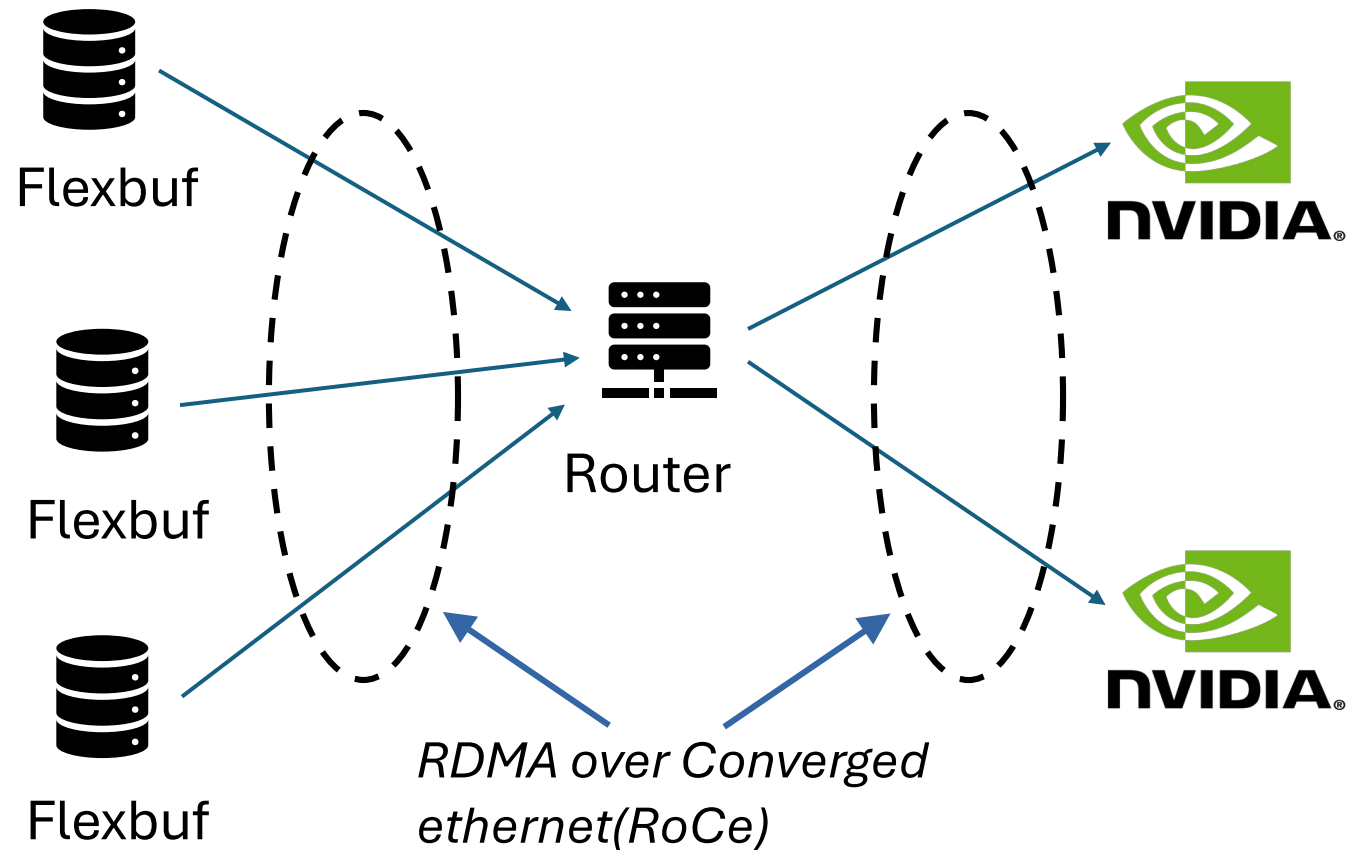
- Created a “ucxwrappers” abstraction
- Can use RDMA directly into GPU memory
 - Even works for “localhost”; using ConnectX NIC as a DMA engine!
- Correlator process consumes almost no CPU cycles!
- Reader process doesn't need CUDA



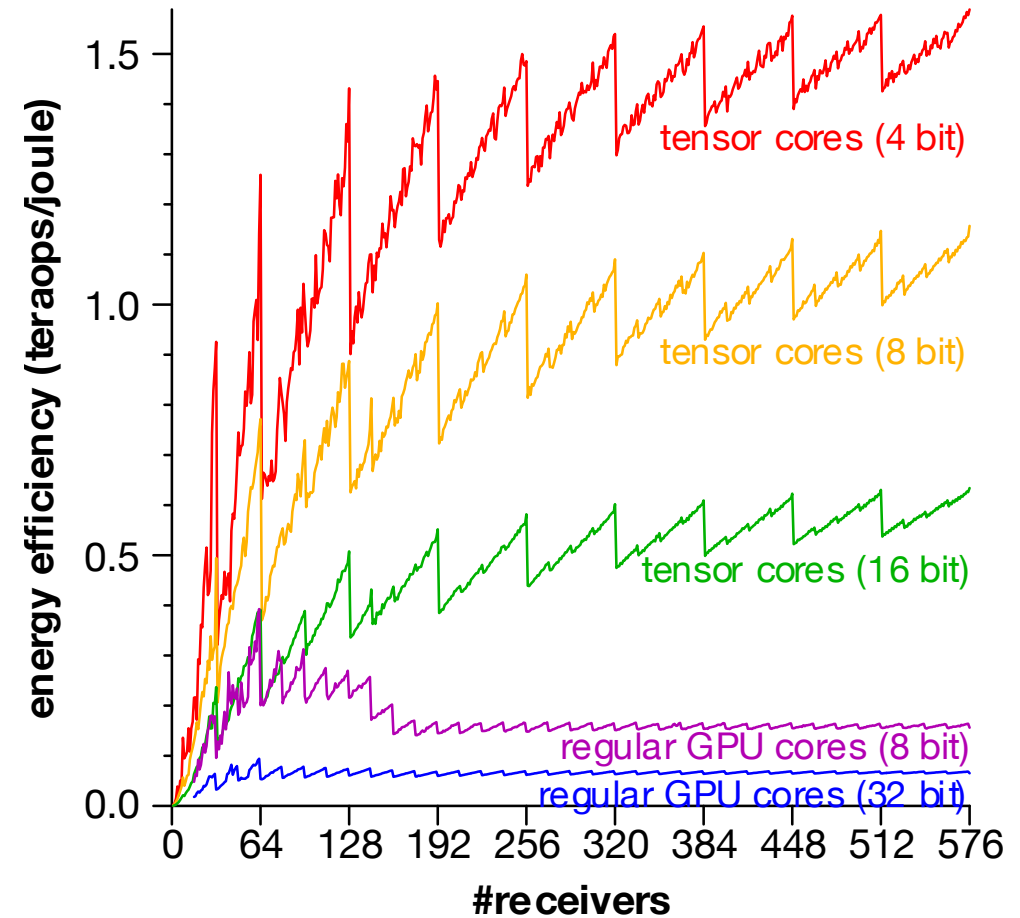
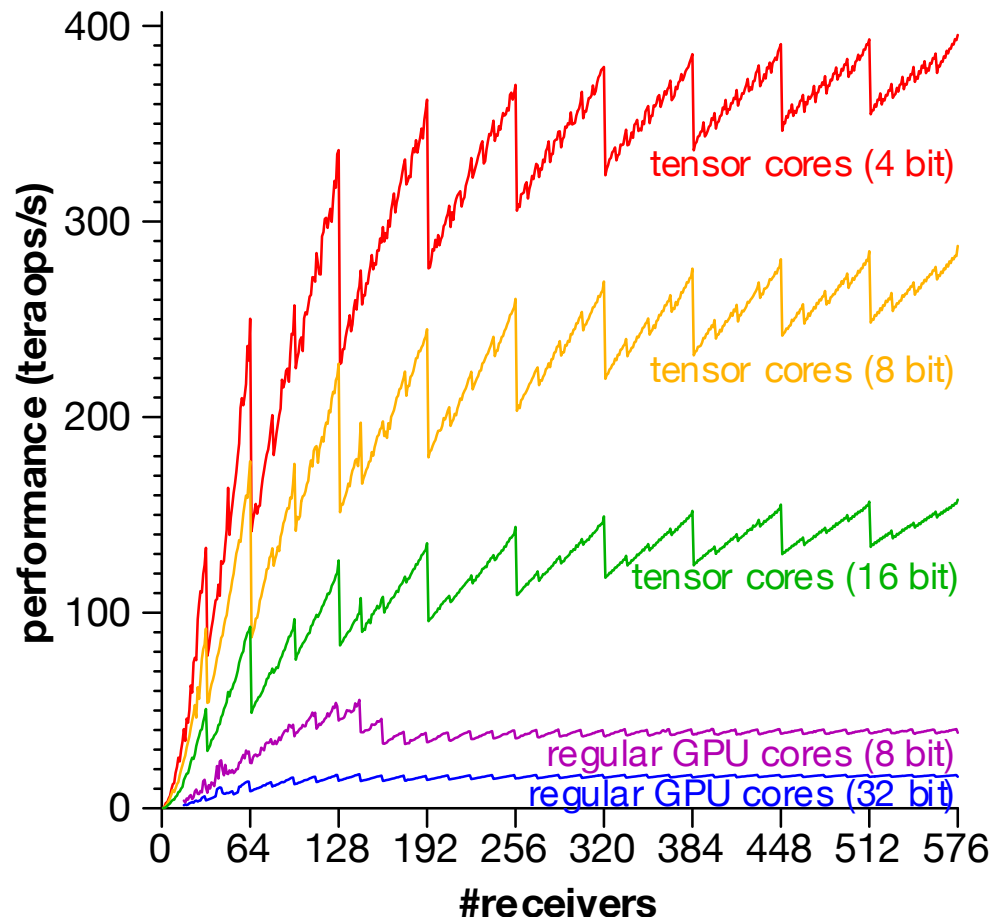
*Unified Communication X
(UCX) framework for HPC*
<https://openucx.org/>

Architecture

- Reader processes on Flexbufs
- VDIF data decoded and corner-turned on Flexbuf
- Coarse integer delay correction is also applied on Flexbuf
- Correlation is parallelised in time and frequency
- Data is streamed directly into GPU memory using RDMA
- RDMA over Converged Ethernet (RoCE) through the ConnectX NICs

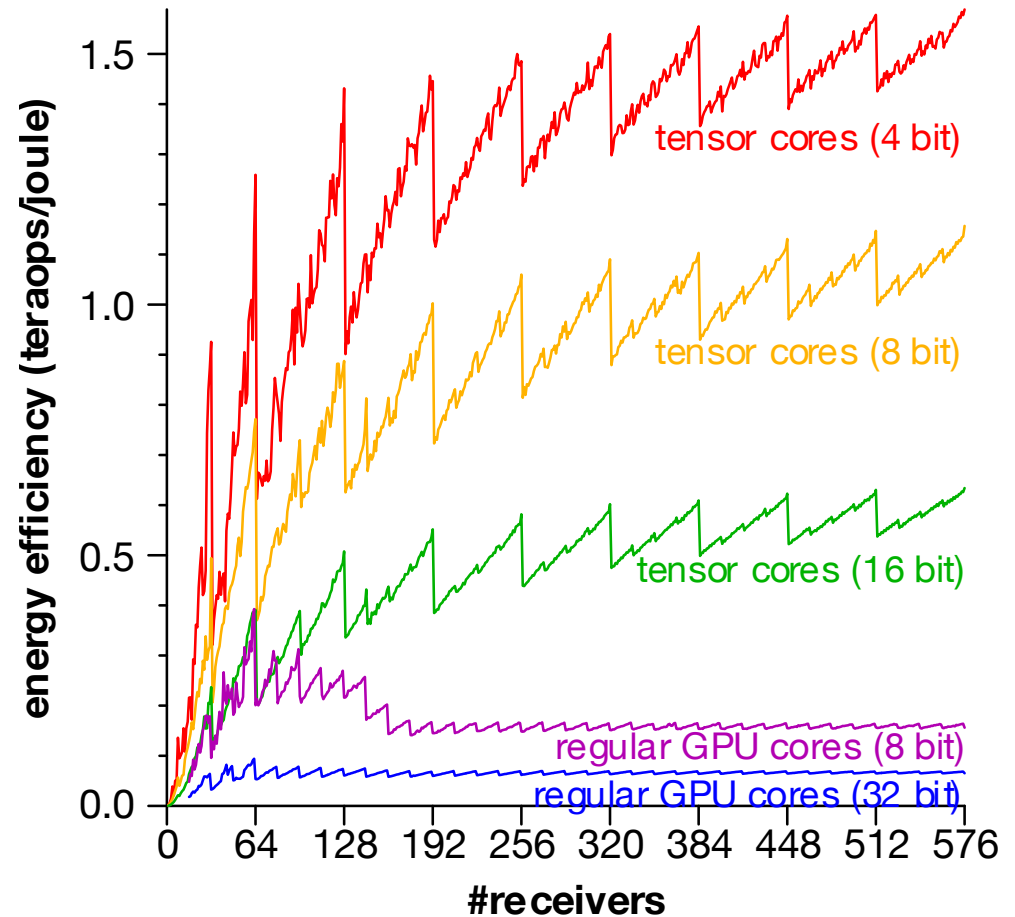
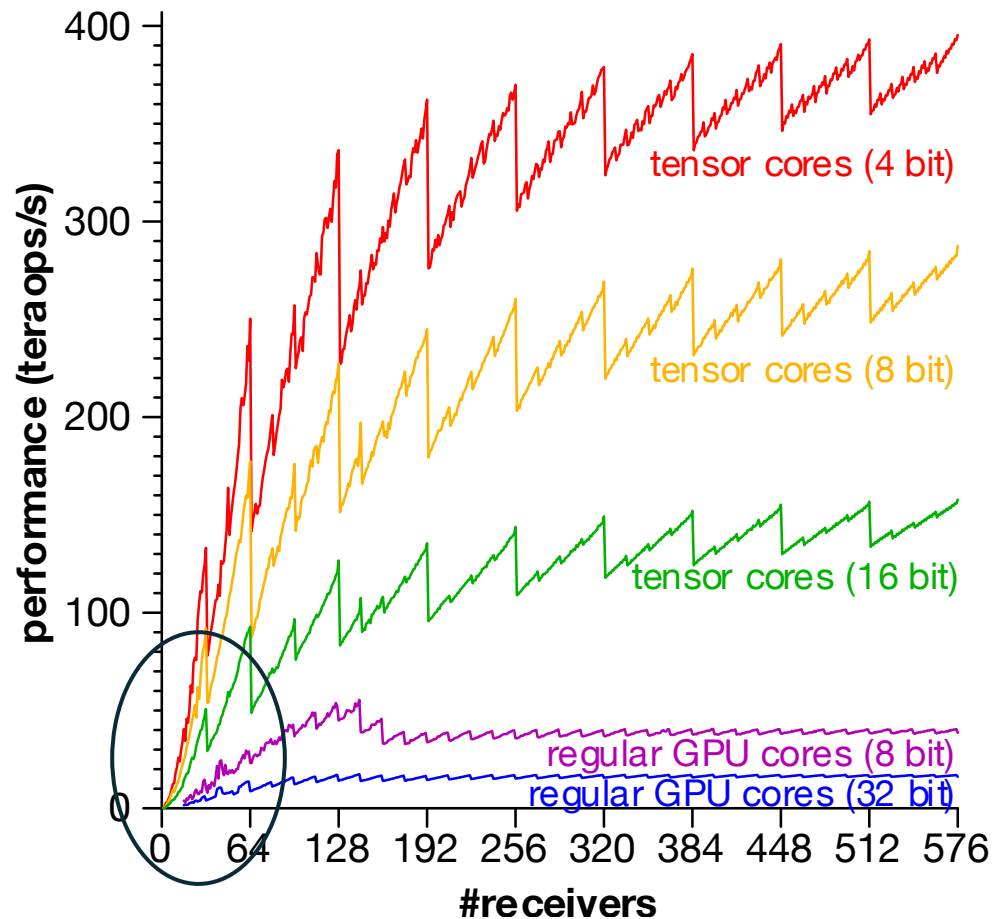


Tensor Core correlator



- John Romein: <https://git.astron.nl/RD/tensor-core-correlator>

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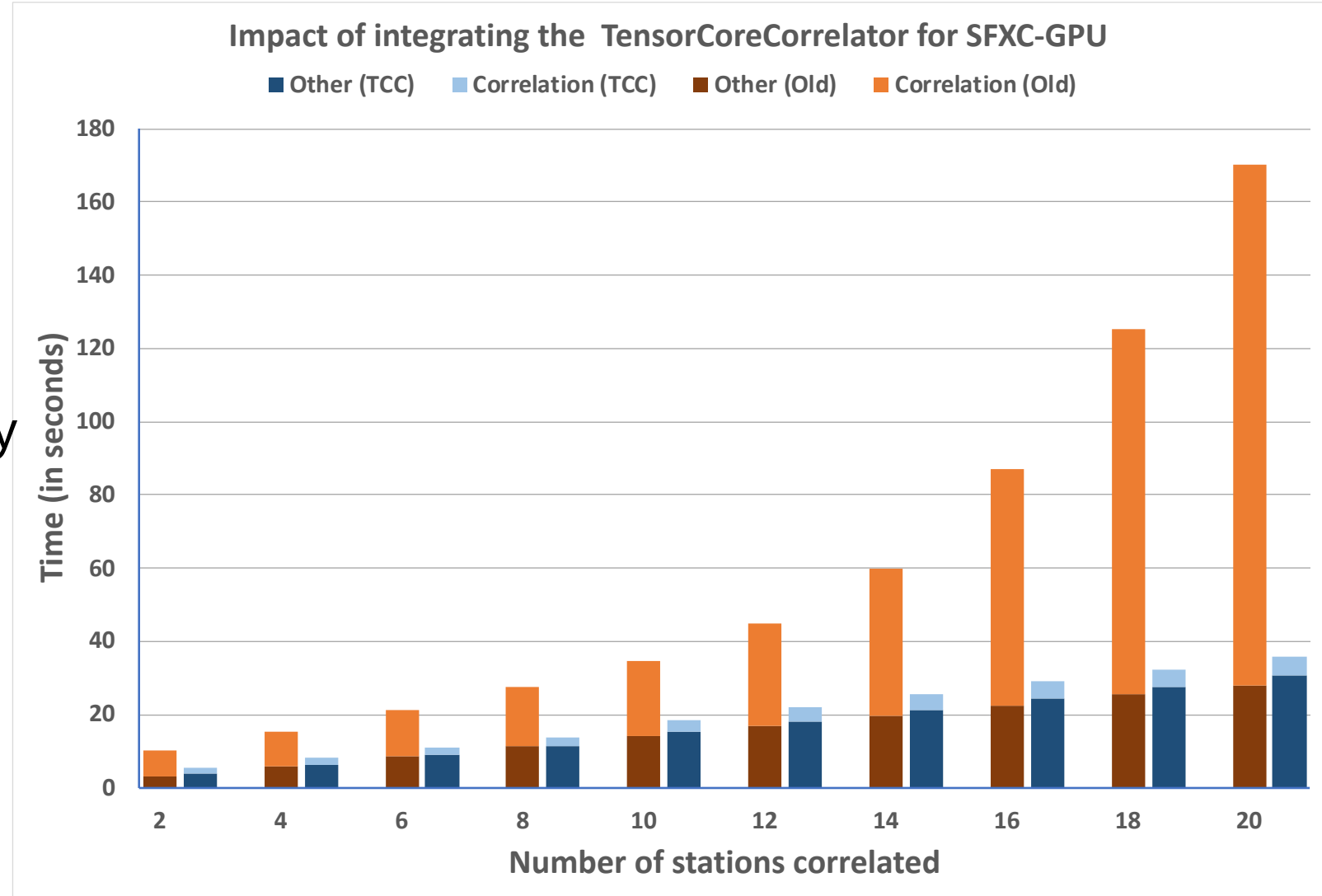


Significant speedup even in VLBI case

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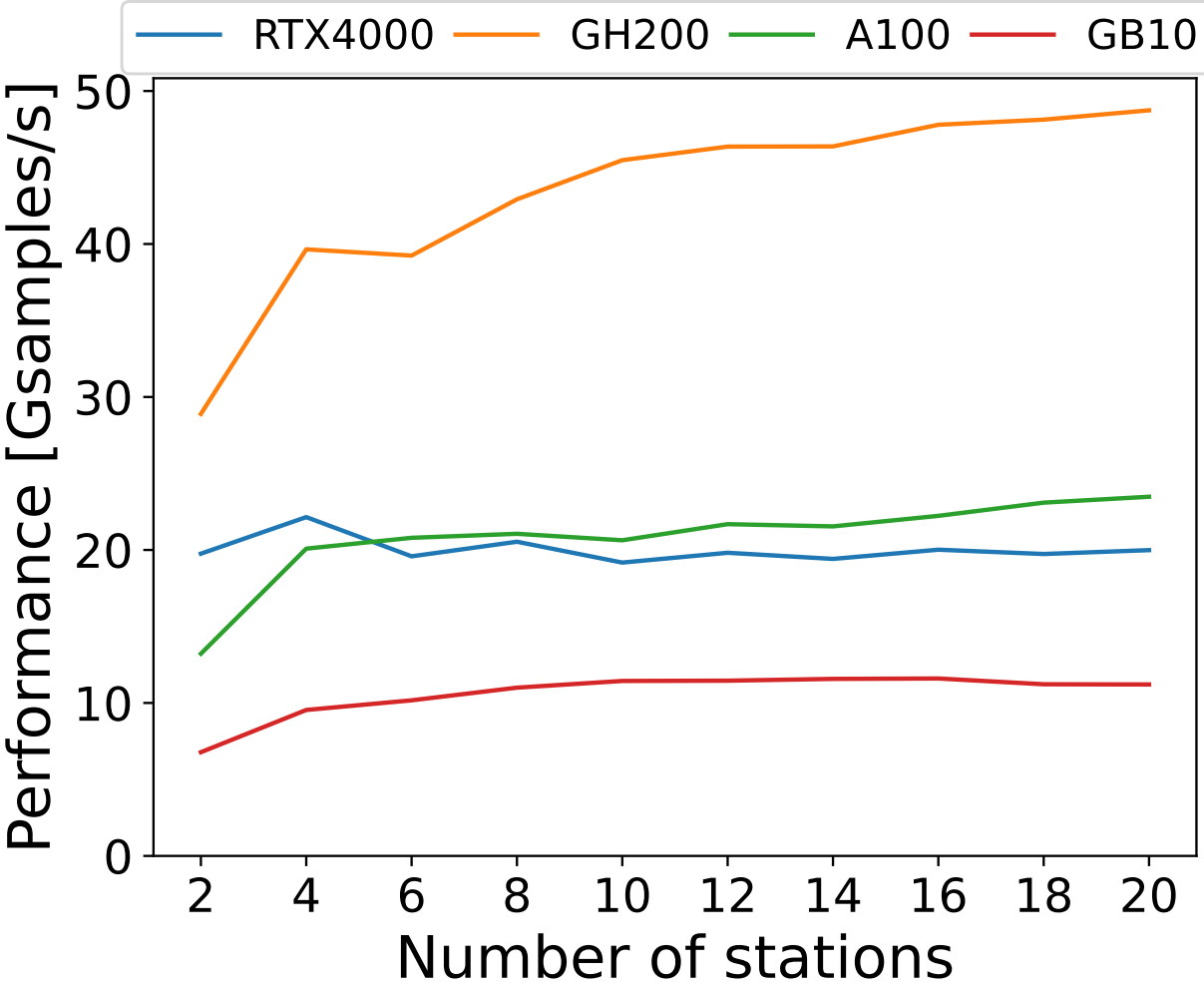
Effect of TCC on SFXGPU

- Old correlation kernel used regular GPU cores
- Was based on xGPU developed by NVIDIA
- (cross-)correlation now is only a small part of computational load
- Delay-compensation + FFT now dominate computation



Benchmarks

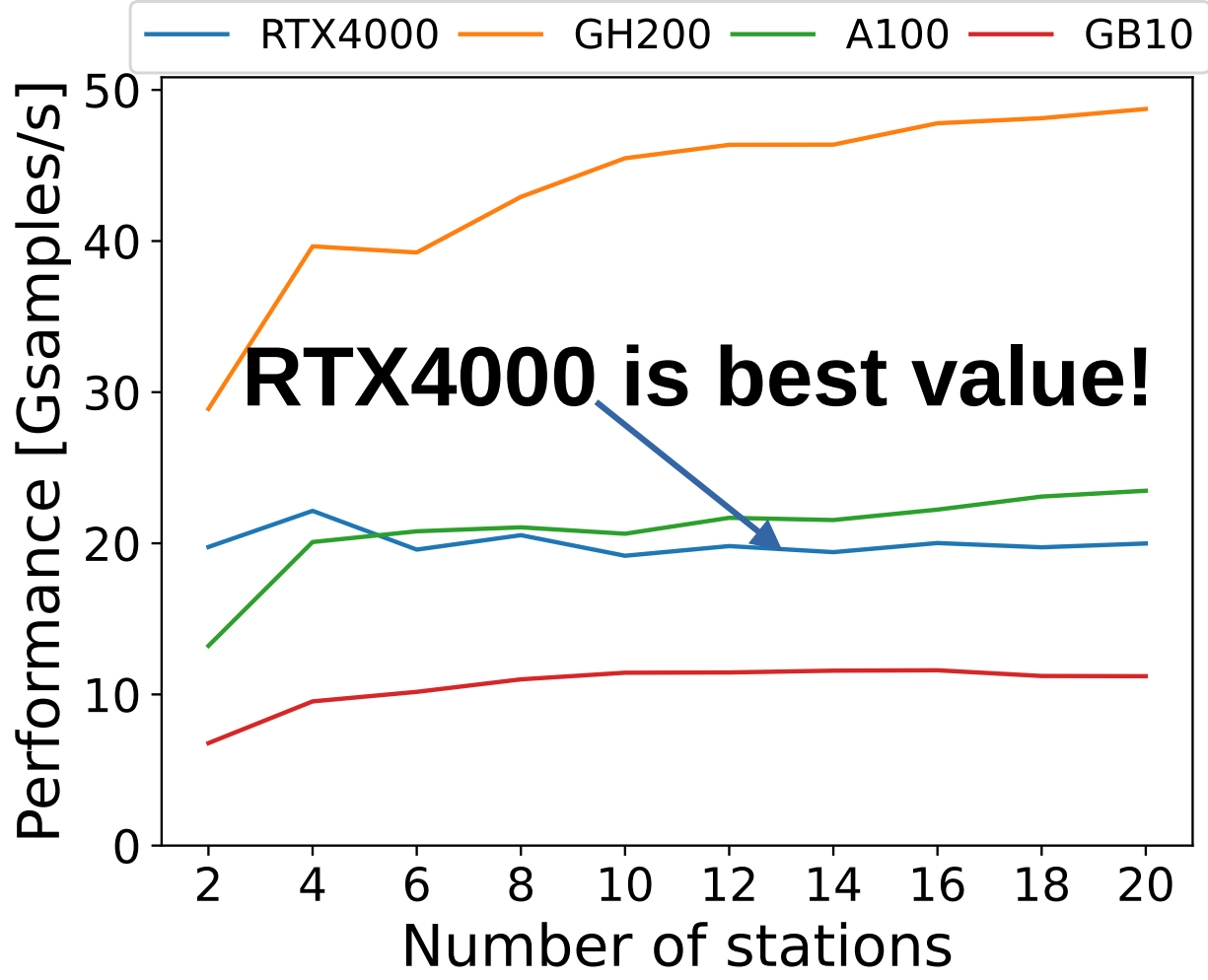
Performance on various GPUs
(256 channels, higher is better)



	RTX PRO 4000	GH200 (Grace Hopper)	A100	GB10 (DGX Spark)
Architecture	Blackwell	Hopper (GPU) + Grace (CPU)	Ampere	Blackwell (GPU) + Grace (CPU)
GPU memory	24 GB GDDR7	96–144 GB HBM3/HBM 3e	80 GB HBM2e	128 GB unified LPDDR5x
Memory bandwidth	~672 GB/s	~4,000 GB/s (HBM) + 900 GB/s NVLink-C2C	~2,039 GB/s	~273 GB/s
TDP	145 W	~900W (system)	400W	~140 W
Price	~€3000	~€30000	~€12500	~€5000

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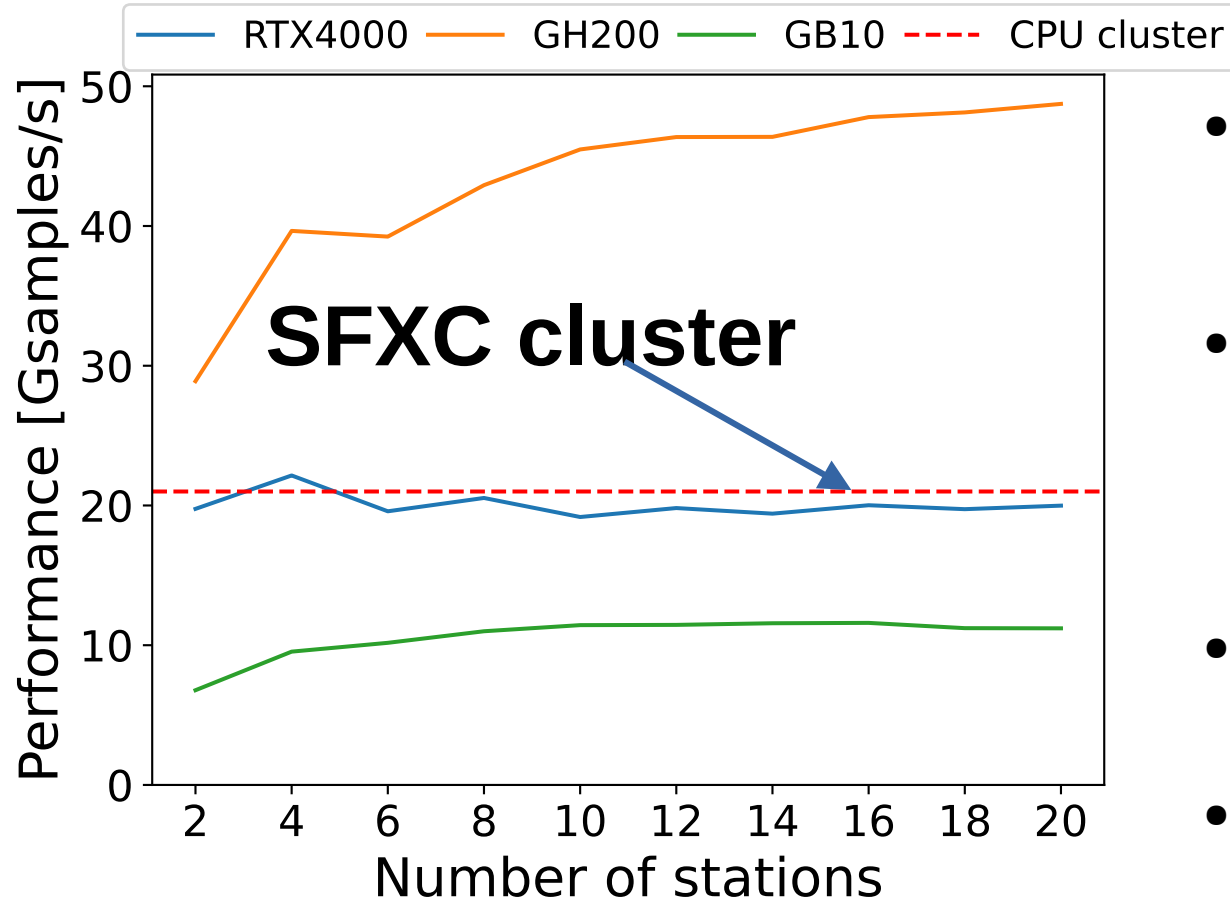
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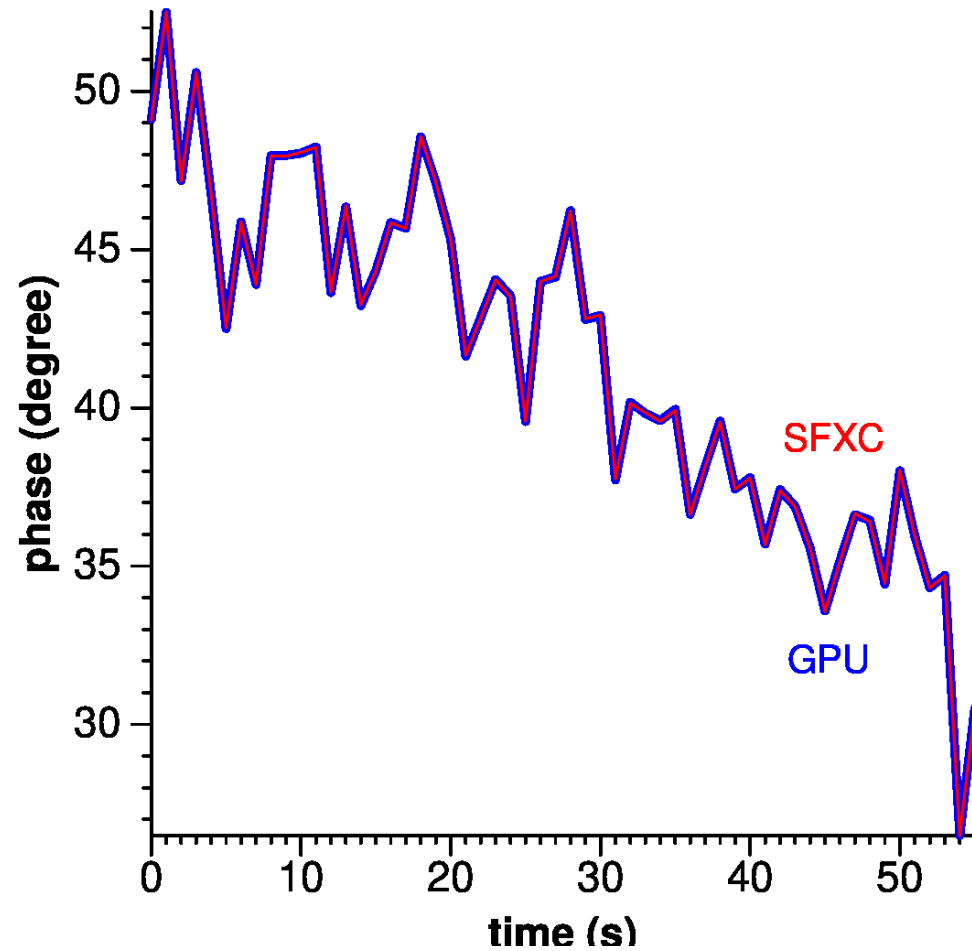
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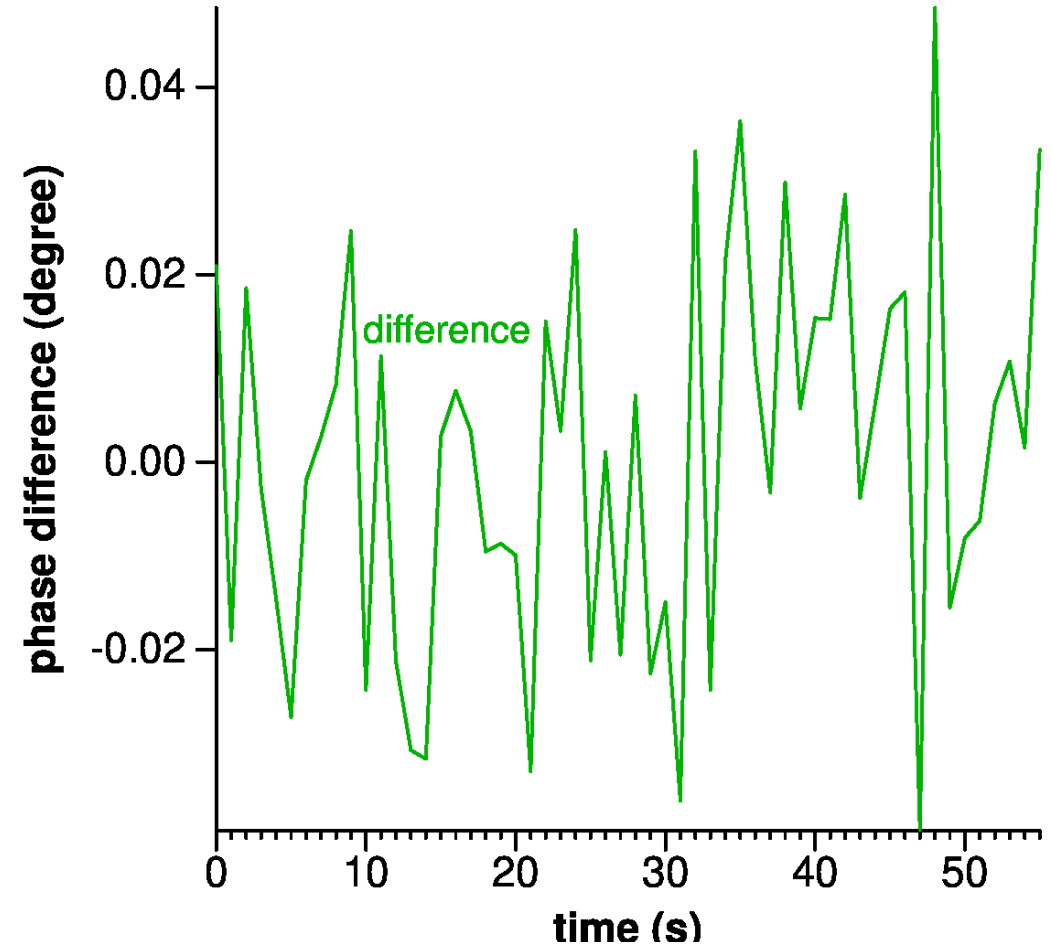


- Current SFXC cluster:
 - 840 Intel Xeon cores (2018)
- Power usage:
 - 3 kW (idle)
 - 7.5 kW (under load)
- Single RTX4000 matches SFXC cluster in performance
- Only 140W per GPU!

Effelsberg – Onsala @ 18cm



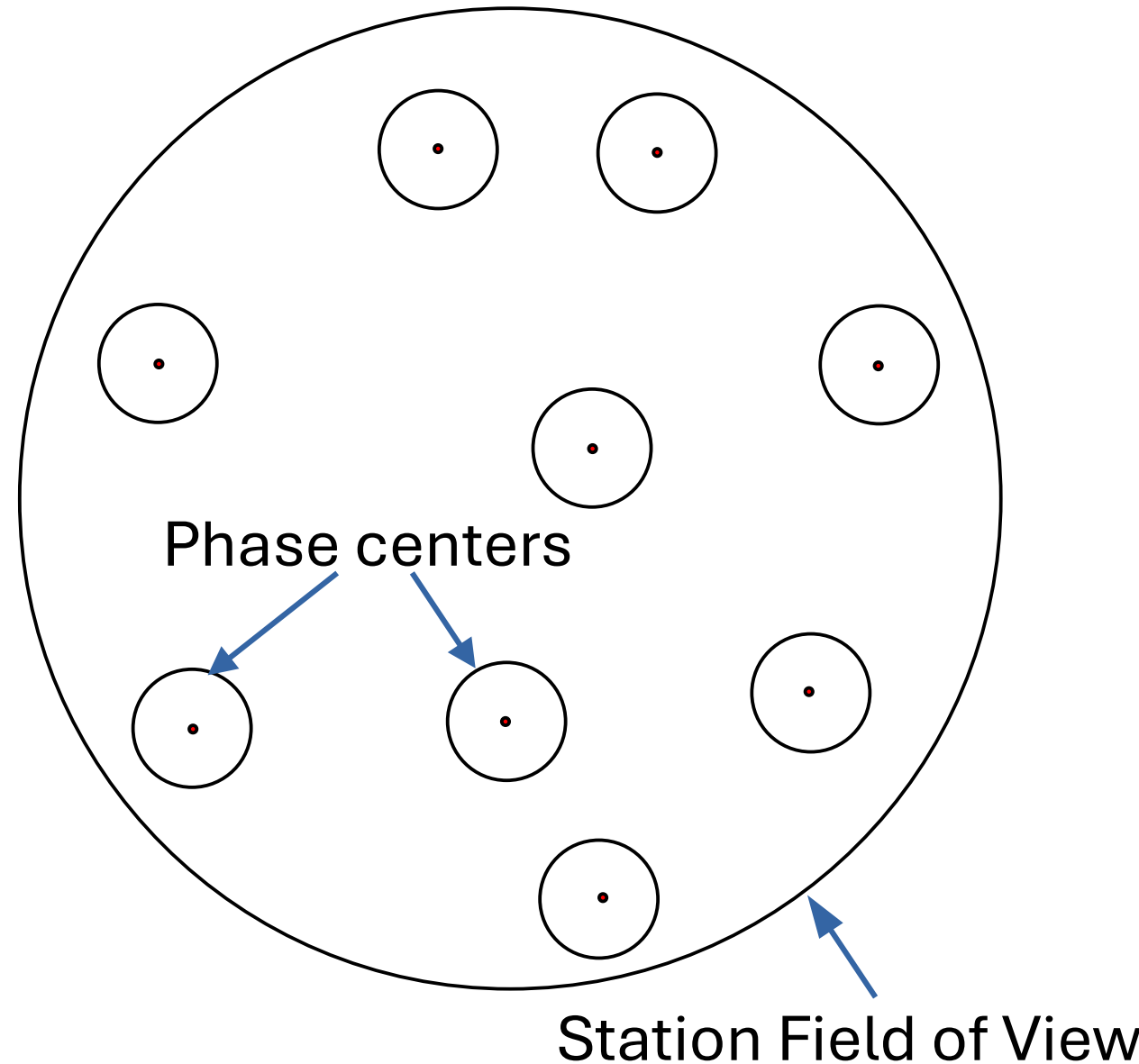
Visibility phase versus time
for both SFXC and SFXGPU.



Phase difference between
SFXC and SFXGPU

Multiple simultaneous phase centres

- It is possible to image the entire field of view of the individual antennas is the array
- **Problem:** Imaging the entire field of view yield very large data sets of mostly empty space
- Typically only a few dozen sources in 10' field of view
- **Solution:** Correlate using high spectral / temporal resolution but only write small narrow field data sets for each source in the field.
- Implementation based on the Tensor Core Correlator

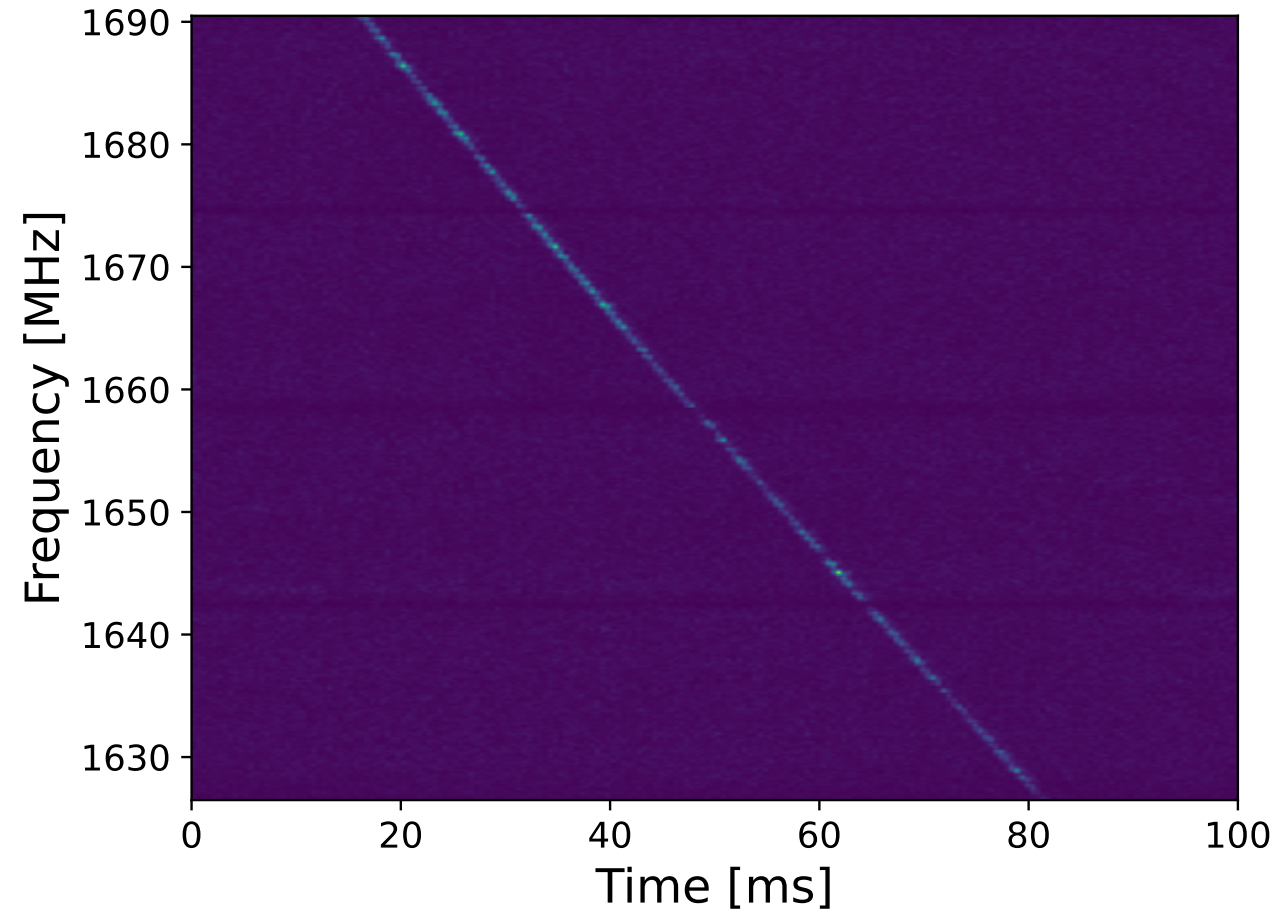


Coherent de-dispersion

- Emission from transients, such as pulsars and fast radio bursts (FRBs) are dispersed by the ISM
- The dispersive delay between two frequencies ν_1, ν_2 [MHz]

$$\Delta t = 4.15 \times 10^{-6} \text{ DM } (\nu_1^{-2} - \nu_2^{-2}) \text{ [ms]},$$

- DM is the dispersion measure
- Apply digital filter to remove dispersive delay from base-band signals
- Computational intensive
 - Very long FFTs, both forward and backward



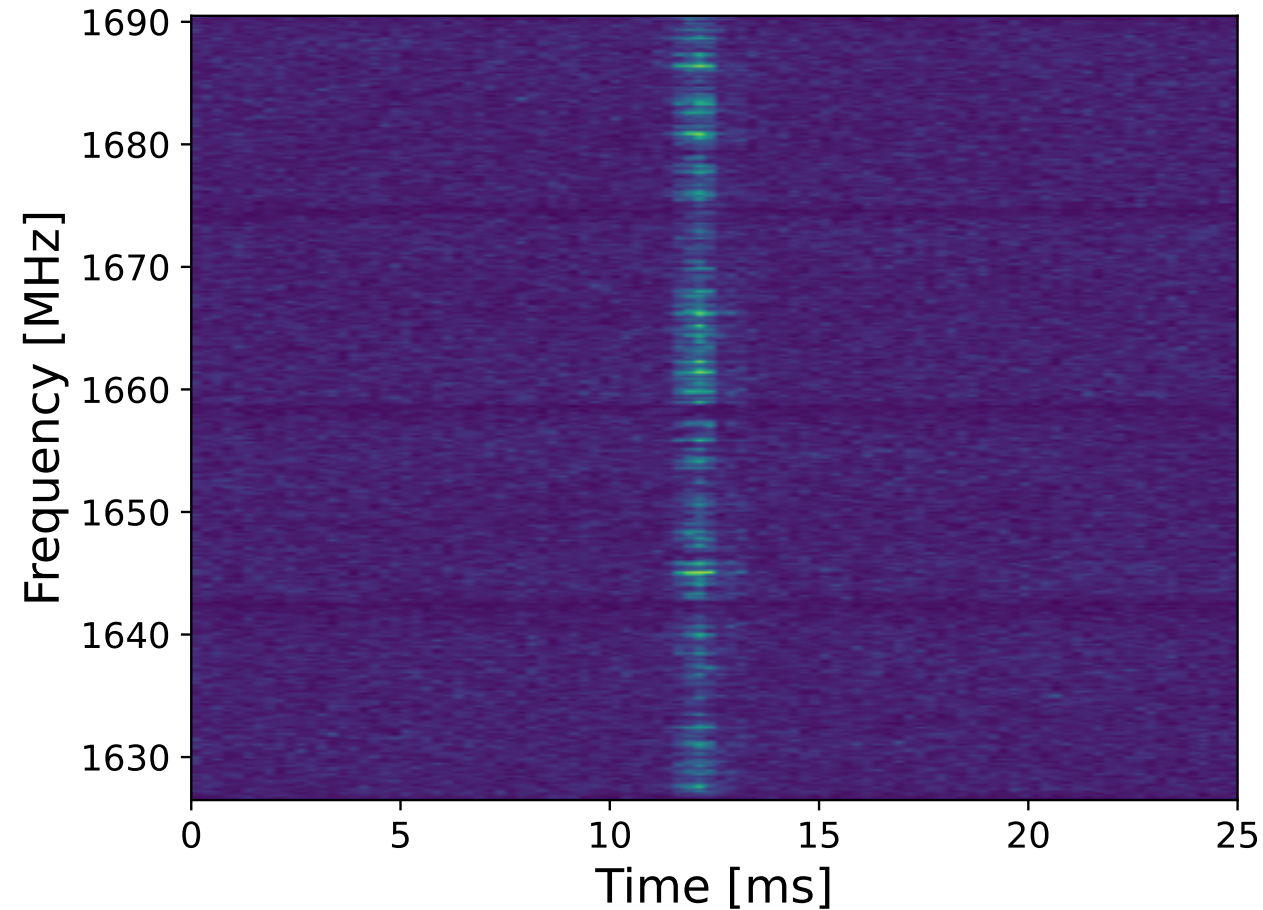
FRB 121102 on Ar-Ef baseline
DM = 561 pc.cm⁻³

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Conclusion and Outlook

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- Most important features are implemented
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