

# LSST Data and Computing

**Nicolas Chotard, Dominique Boutigny**  
**CNRS/IN2P3/ASTERICS**

**WP3 OBELICS Meeting, UCM Madrid**  
**September 14th-15th-16th, 2016**



LSST is an instrument designed to make high precision images of the whole accessible sky in 4-D (x, y, z, t) - 10 year time-lapse movie of the sky

## **Time domain science**

- Novae - Supernovae - GRBs
- Source characterization
- Instantaneous discovery

## **Moving sources**

- Asteroids and comets
- Proper motions of stars

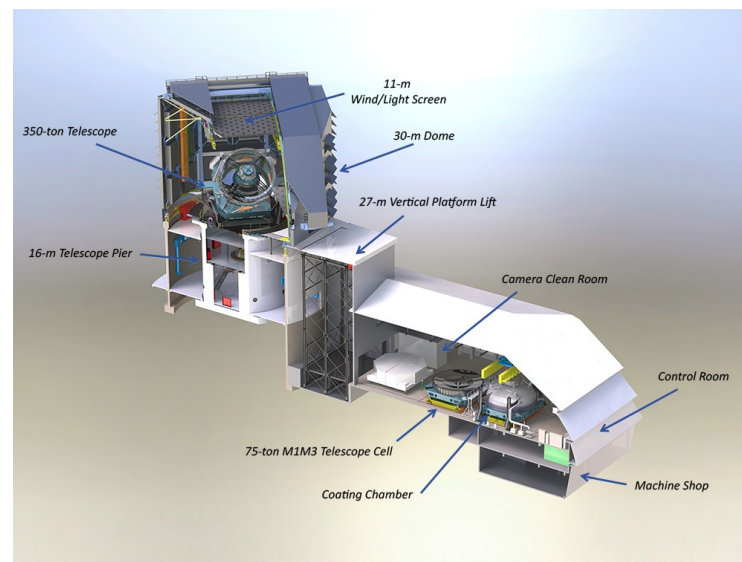
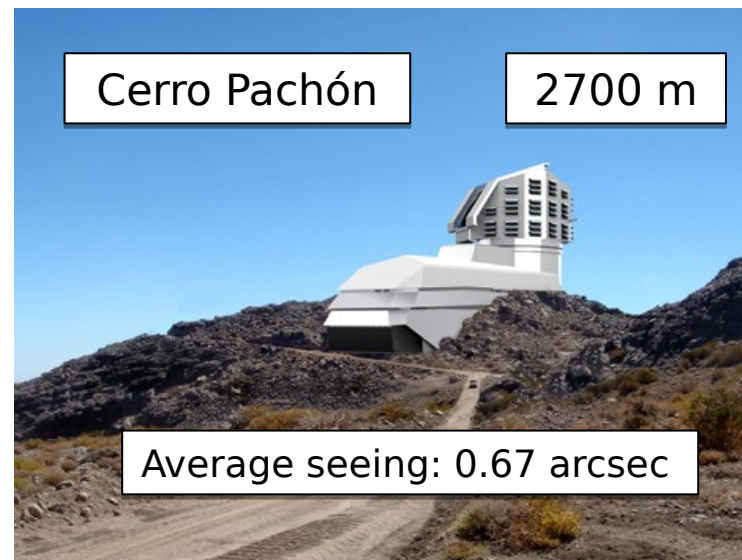
## **Mapping the Milky Way**

- Tidal stream
- Galactic structure
- Complementary to GAIA

## **Dark energy and dark matter**

- Gravitational lensing (strong and weak)
- Evolution of large scale galactic structures
- Trace the nature of dark energy

3 keywords : **Fast - Wide - Deep**





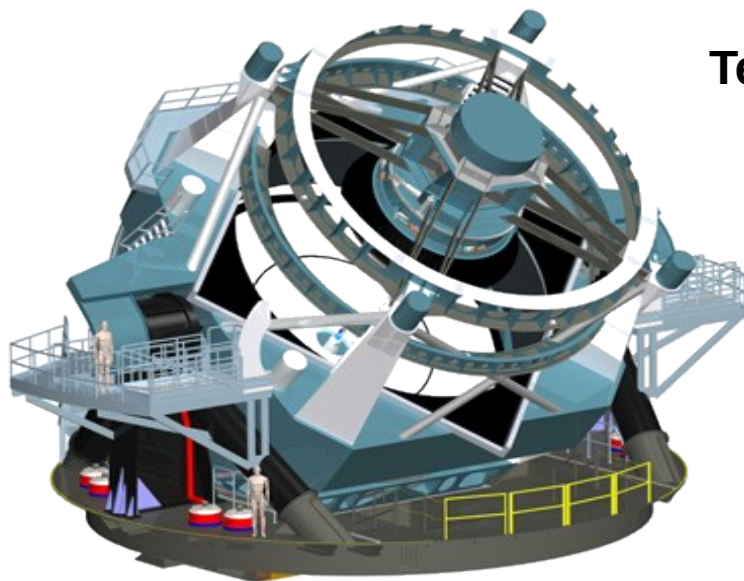
September 9th





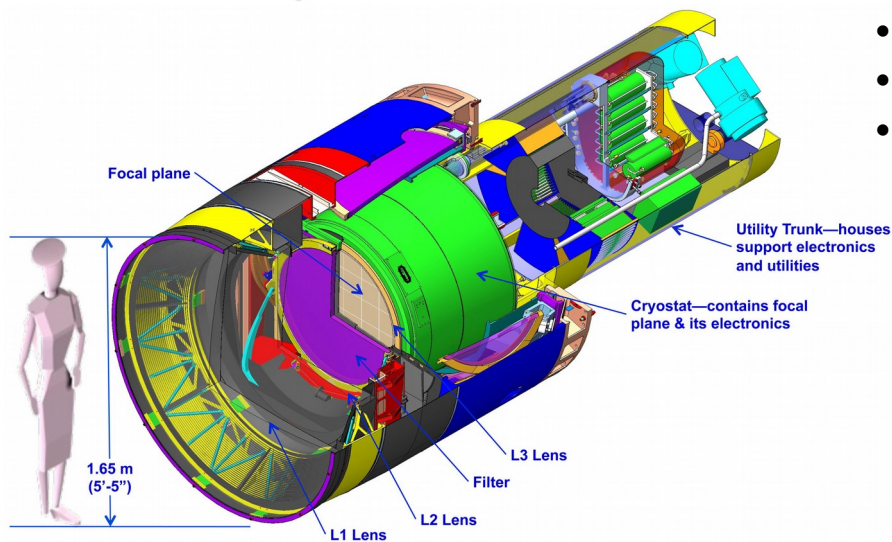
## Telescope

- 8.4 m (6.7m effective)
- Fully corrected (sphericity, coma, astigmatism)
- A 350 tons mobile structure



## Camera

- 3.2 billion pixels @ 0.2 arcsecond / pixel
- 21 rafts
- 9 CCD / raft



Camera 3/4 Section



- **Timing**

- 1 visit = 2 successive exposures
  - CCD Exposure: 15 s
  - CCD readout: 2 s
  - Shutter positioning : 1 + 1 s
  - Telescope slewing : 5 s (median)
- **Total / visit: 39 s (median) - 44 s (mean)**
- The whole visible sky is imaged every 3 days
- Alerts on transient events will be available world-wide within 60 sec
  - ~10 millions / night

- **Size**

- **~15 TB of raw data / night**
- Repeat this during 10 years starting in 2022
- **Final image collection : 0.5 Exabytes**

- **Field**

- **Deep Wide Survey: 18,000 deg<sup>2</sup>**
- Deep Drilling Survey: ~30 selected fields - 300 deg<sup>2</sup> : ~1 hour /night
- Most of the sky will be covered over 800 times with 2 × 15 sec visits
- Exposure of the same field are stacked together (co-addition)
  - reach magnitude 27.5

## LSST DATA CENTERS



Summit (Pachón) to Base (La Serena) = 2 x 100 Gbps  
 Base(La Serena) to Archive (NCSA) = 2 x 40 Gbps

- Collaboration between IN2P3 and LSST / NCSA
  - 50 % of the data processing → annual catalogs production
  - Integral copy of all data at CC-IN2P3
  - A 20 Gb/s dedicated link already deployed
- Explore standard transport protocols for bulk file transfer over high-latency network links
  - Goal: understand what will be possible to use for transferring data between CC-IN2P3 and NCSA for the 2020-2030 era
  - Test HTTP2 as an underlying transport protocol
- Test Object Store technologies to determine if they are suitable for LSST use cases
  - OpenStack Swift and CEPH



- Goal: develop a full modular, efficient and versatile image analysis framework
- LSST software stack is largely written from scratch, in Python, unless computational demands require the use of C++
  - **C++:**
    - Computationally intensive code
    - Made available to Python via SWIG
  - **Python:**
    - All high-level code
    - Prefer Python to C++ unless performance demands otherwise
  - All the code is, and will stay open source
- But we do not re-invent the wheel
  - Use of standards (fits,...)
  - Implantation of algorithms developed by others (SDSS,...)
  - Interfaced with existin software (psfex, GalSim,...)
- Modularity
  - Virtually everything is a Python module
  - ~60 packages ([git repositories](#), ~corresponding to python packages)
- Designed to support several instruments: LSST - SDSS - HSC - CFHT - DES

## Three data product levels

- A stream of ~10 million time-domain events per night, detected and transmitted to event distribution networks within 60 seconds of observation.
- A catalog of orbits for ~6 million bodies in the Solar System.
- A catalog of ~37 billion objects (20B galaxies, 17B stars), ~7 trillion observations ("sources"), and ~30 trillion measurements ("forced sources"), produced annually, accessible through online databases.
- Deep co-added images.
- Services and computing resources at the Data Access Centers to enable user-specified custom processing and analysis.
- Software and APIs enabling development of analysis codes.

Level 1

Nightly data product

Level 2

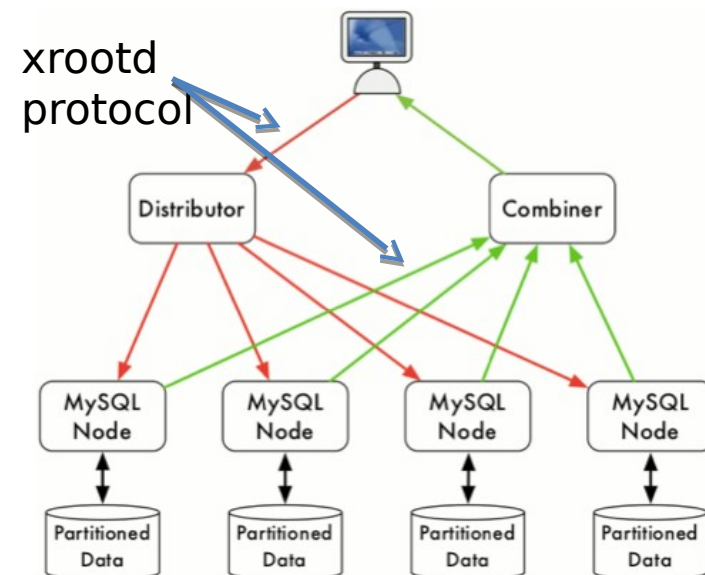
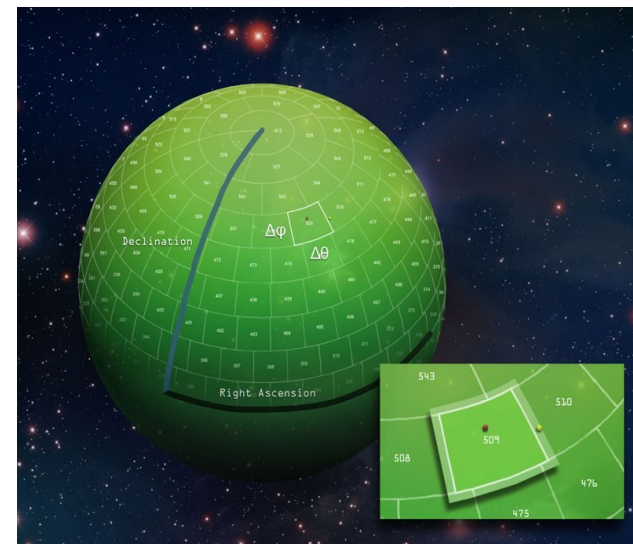
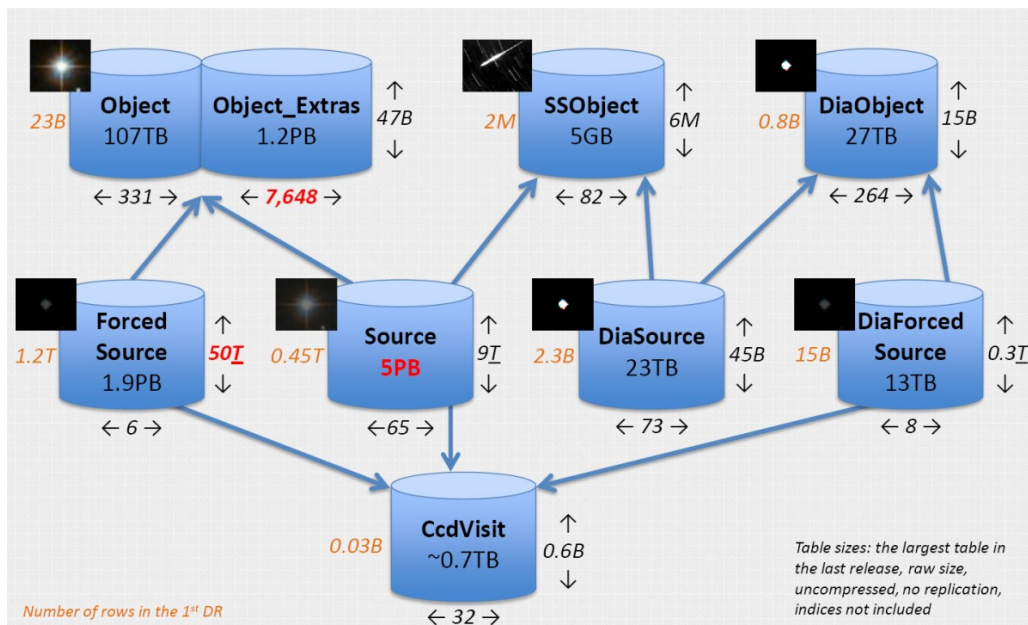
Annual data product

Level 3

User-created data product

The challenge is to design an SQL database system able to store trillions of objects while keeping a reasonable access time

**Qserv** : developed at SLAC + IPAC  
 Design optimized for astronomical queries

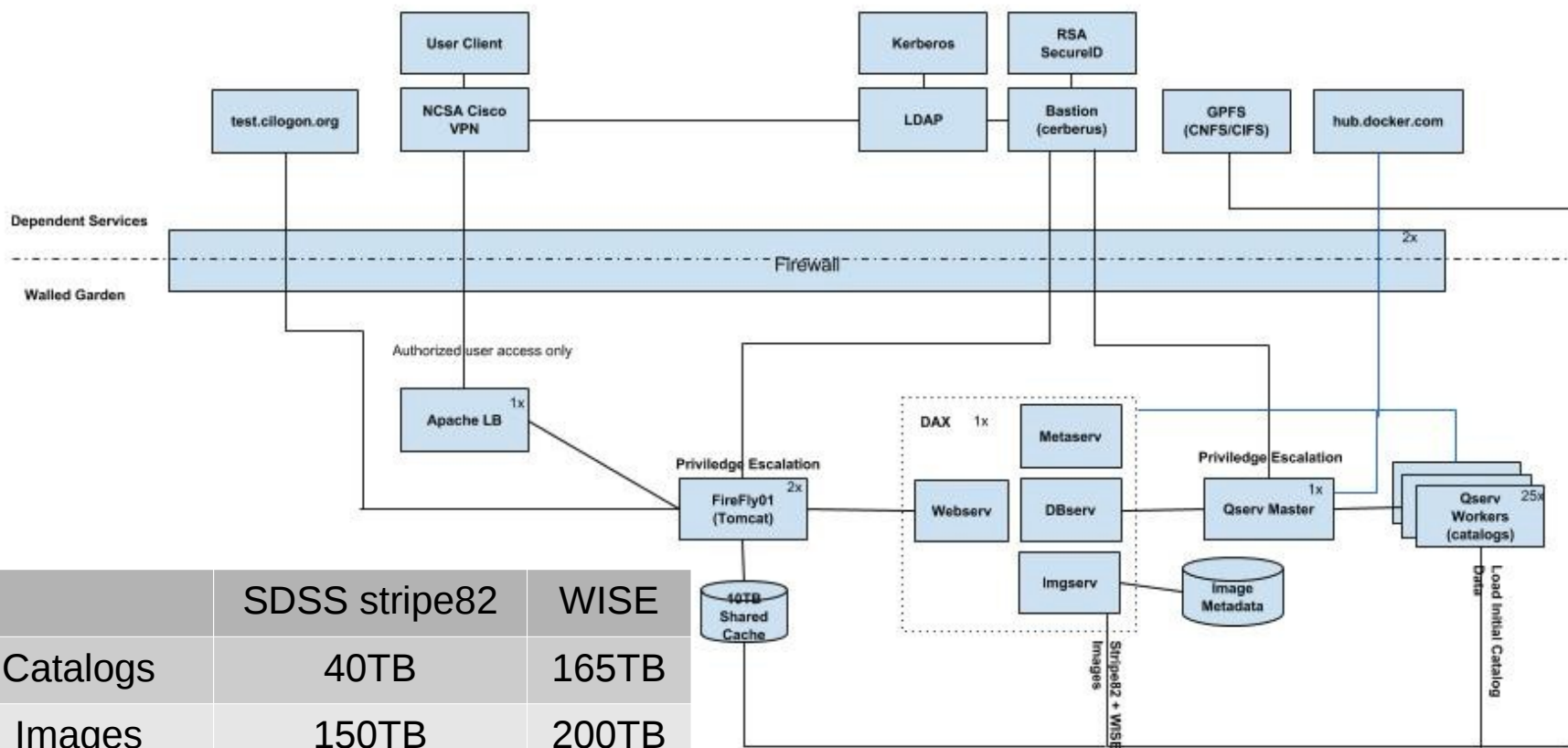


Massively parallel – distributed – fault tolerant **relational database**



- Thanks to a partnership with Dell we have deployed a Qserv test bench
- The only test bench currently available in LSST for large scale tests
  - 50 nodes - 400 cores
  - 800 GB memory
  - 500 TB disk storage
  - Non volatile memory devices to front end machines
  - Currently include 35TB of real data
  - Next step: 120TB from the Wide-field Infrared Survey Explorer (WISE)
- Test bench currently replicated at NCSA
- An other computer is lend by DELL for local Qserv tests at CC-IN2P3
  - CPU: 2 x Intel Xeon E5-2680 v4 (2.4 GHz, 35 MB cache, 14 cores, 28 threads)
  - RAM: 512 GB
  - Persistent storage
    - Data: NVMe (2 TB); SSD (3.2 TB); HDD (10 TB)
    - System: 0.6 TB )
  - Network: 10 Gbps
  - Operating system: CentOS 7, /sps/lsst, etc.

- Currently under preparation at NCSA
- Will be duplicated at CC-IN2P3 in a few month
- Also used to test the DB replication between the two distant sites



	SDSS stripe82	WISE
Catalogs	40TB	165TB
Images	150TB	200TB
Metaserv DB	< 1TB	< 1TB

## Main technical implications

- Tests of the LSST stack
  - Validation on real data
  - Integration of new algorithms
  - Scientific validation of the data base (Qserv) functionality
- bridge between software and science  
→ **all these activities have been started quite recently**



- LSST is going to push astronomical data to an unprecedented scale
- Project is approved and will operate in 2022 - 2031
  - Site, camera, mirror, filters all under construction
- LSST is building a new open source, multi-instrument, software stack
  - C++ and Python
  - Currently under development and tested against real data
- A highly distributed / fault tolerant SQL database system to store and give access to the catalogs is being developed: **Qserv**
  - Test on real data (clusters in CFHT images) will start soon
- Yearly Data Release Processing will be split between NCSA and CC-IN2P3
- **Still need to precise how to articulate the LSST use case within OBELICS work packages**

## Observation Properties:

The standard visit = 2 x 15 sec. exposures

Exposure sequence = 1 s open shutter + 14 s dwell + 1 s close shutter = 16 s

Visit sequence = 16 s exp + 2 s readout + 16 sec exp = 34 s

(second readout concurrent with slew)

Median slew time between visits = 5 s

Average slew time between visits = 12 s

Visits per night = "about a 1000"

Calibration exposures = 450/day

Data collected per 24 hr period = "about 15 TB"

## Network bandwidths:

Summit (Pachón) to Base (La Serena) = 2 x 100 Gbps

Base(La Serena) to Archive (NCSA) = 2 x 40 Gbps

## Alert Production:

Real-time alert latency = 60 seconds

Average number of alerts per night = "about 10 million"

## Data and compute sizes:

Final image collection (DR11) = 0.5 Exabytes

Final database size (DR11) = 15 PB

Final disk storage = 0.4 Exabytes

Peak number of nodes = 1750 nodes

Peak compute power in LSST data centers = 1.8 PFLOPS

## Data Releases:

Number of Data Releases = 11

DR1 release = Date of Operations Start + 12 months

Estimated numbers for DR-1 release

Objects = 18 billion

Sources = 350 billion (single epoch)

Forced Sources = 0.75 trillion

Estimated numbers for DR-11

Objects = 37 billion

Sources = 7 trillion (single epoch)

Forced Sources = 30 trillion

Visits observed = 2.75 million

Images collected = 5.5 million

[More on the LSST webpage](#)