

GW detectors and related institutions

- Einstein Telescope (3rd generation detector)
 - FP7 design study (2011).
 - ASPERA roadmap. Candidate ESFRI
- Advanced Virgo "pathfinder" (2nd generation)
- European Gravitational Observatory, EGO
 - CNRS-INFN consortium with other partners
 - Manages Virgo site (Italy) and hosts ET coordination





Status of Advanced Virgo

- Currently upgrading: x 10 sensitivity improvement over 1^{st} generation (2007-12) $\rightarrow x$ 1000 in the event rate
 - Observability horizon for binary neutron stars : 140 Mpc
 - Current BNS event rate estimates: few to tenth events/yr
- Plan: First science data in 2016 jointly with US based LIGO
- Opportunities for **multimessenger astrophysics**
 - Search for electromagnetic counterpart (i.e., GRB afterglow, flaring object, kilonova, ...)
 - Extensive electromagnetic follow-up program including LOFAR, HESS, CTA, ...
 - Interoperability between observatories being laid out





GW related-challenges

- Low-latency analysis of distant detectors
 - Should allow background rejection "glitches"
- Alert generation within minutes
 - To allow the observation of prompt emission and short afterglows
- GW error is large by astronomical standards
 - Few 100 sq degrees Needle in the haystack: many galaxies and many potential unrelated transient sources
 - Prioritize regions to follow-up





Multimessenger astrophysics with GWs

- Connection with ASTERICS topics and WPs
 - Low-latency alert generation from GW data (WP5)
 - Optimization of follow-up strategy (in large GW error box)
 Cross-correlation with catalogs of nearby galaxies (WP4)
 - Cross-matching with other astronomical transients (WP3 & WP4)
 Rejection of unrelated transients (variable stars, AGN, ...)
 - Robust assessment for coincidence significance (WP3)
- Expectations from ASTERICS
 - Expertise and knowledge sharing (Virtual observatory, other ESFRI)
 - New contributors





GW observations and the VO

- Access model for GW data in the next decade
 - Assessing GW event significance is difficult. This is the reason with we operate currently under a close-data model
 - However, we have a **plan for public release**
 - After first 4 alerts, https://dcc.ligo.org/LIGO-M1300550
 - LIGO open science center, https://losc.ligo.org
- Expectation for ASTERICS
 - Prepare/support future public release to the VO